

Search for New QCD Phenomena at the **STAR* Experiment at RHIC**

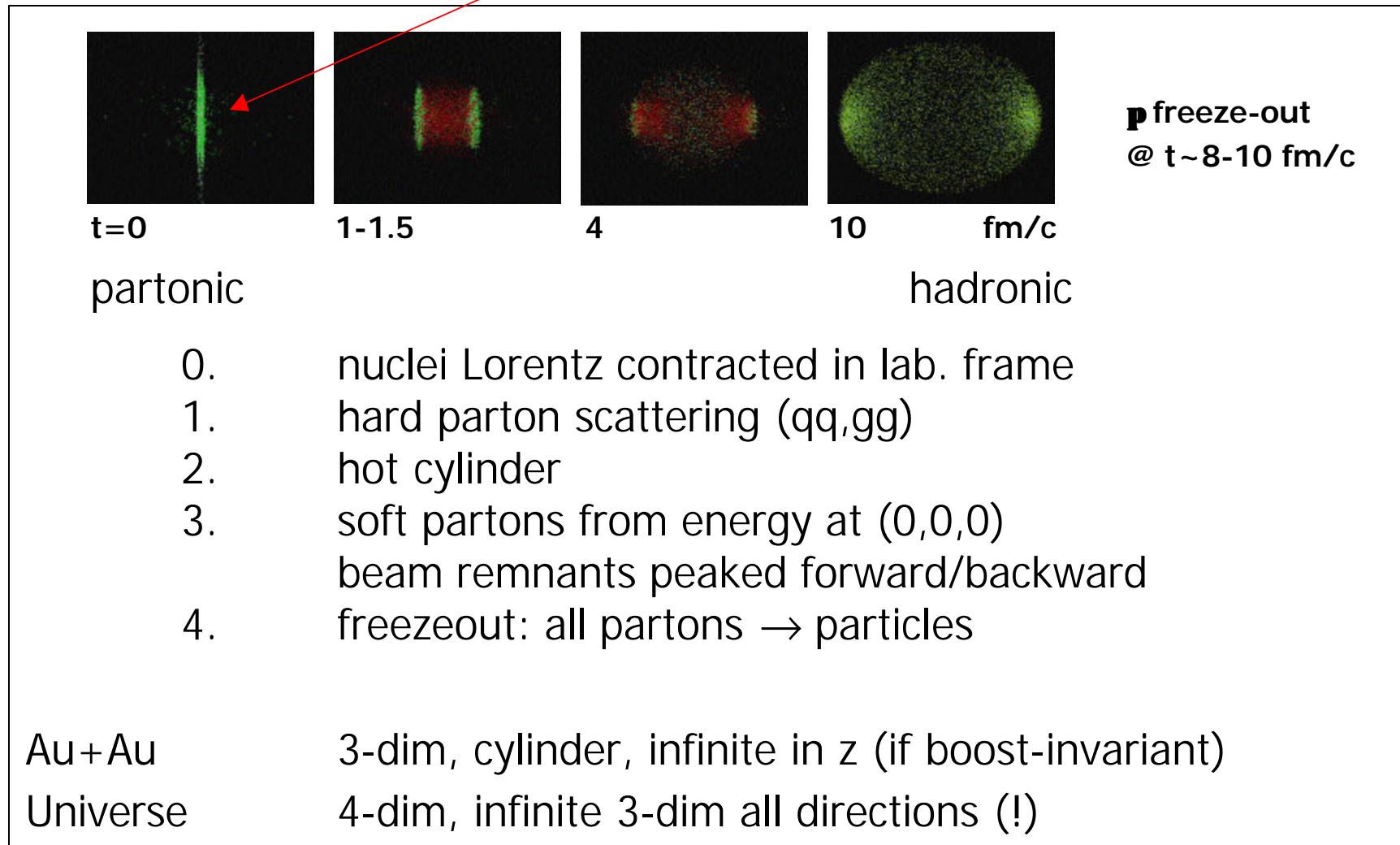
*** (Solenoidal Tracker at RHIC)**

Jens Sören Lange (Frankfurt/RIKEN-BNL)
on behalf of the STAR Collaboration

4th Tropical Workshop on Particle Physics and Cosmology
Cairns, Australia, June 9, 2003

1 Au+Au Collision

QM ($\Delta x \Delta p < h$) formally induces $v > c$



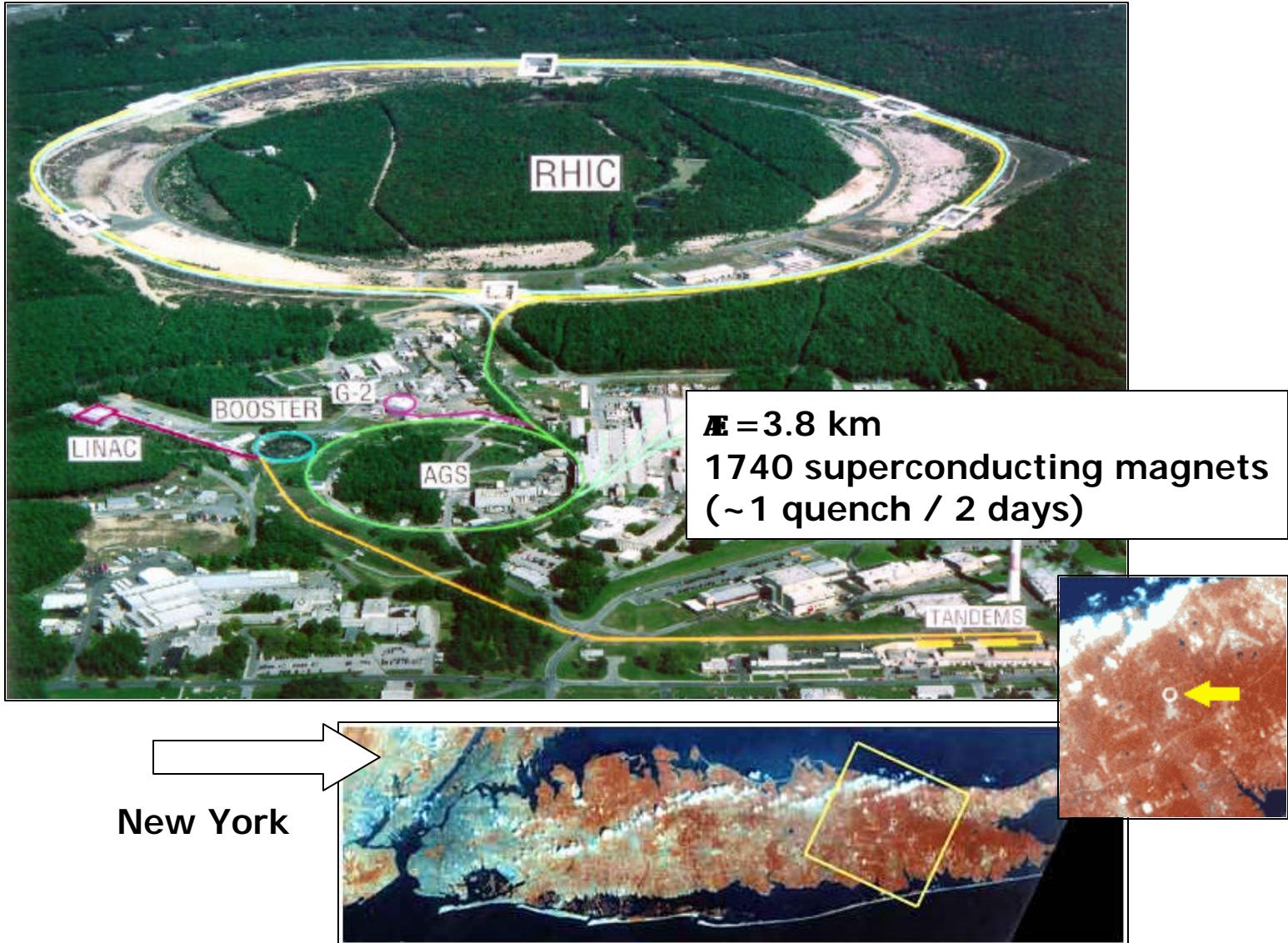
What we want to know ?

- Is there order in this mess ?
(rules of particle production)
- Temperature ?
- Size ?
- Density ?
- How can we produce Anti-He ?
- The little and the big bang -
are the rules the same ?

Au+Au

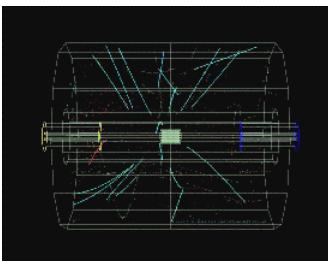
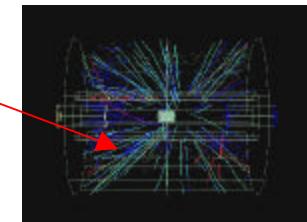
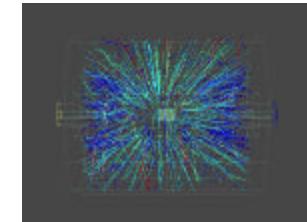
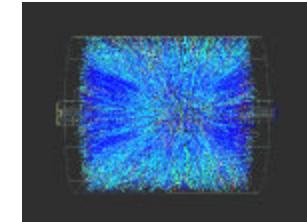
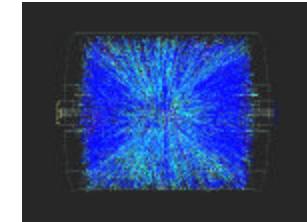
2000: $\sqrt{s_{NN}}=130$ GeV

2001: $\sqrt{s_{NN}}=200$ GeV



Data set.

- Au+Au 130 GeV 0.7 Mio
- Au+Au 200 GeV 3.2 Mio
- Au+Au 19.6 GeV ~20k
- d+Au 200 GeV 35 Mio.
- pp
 - un-polarized
 - vertical pol. 391/nb
 - longitudinal pol.
(spin flip snake) 373/nb
- Level-3 trigger, rare probes
- EMC jet trigger



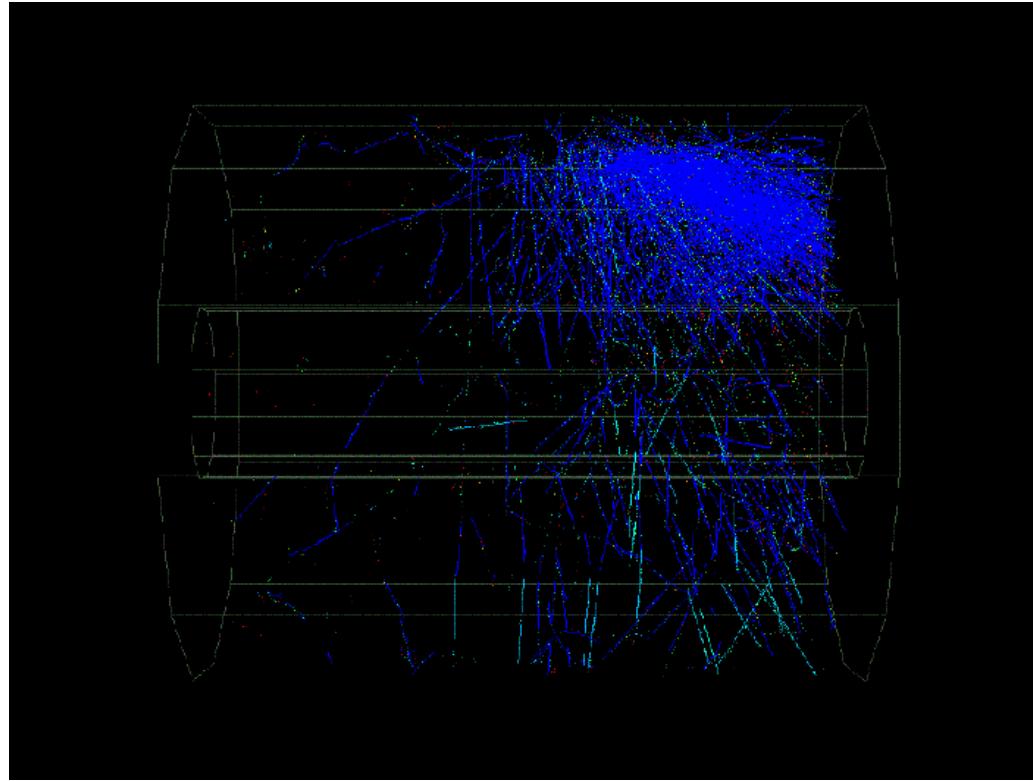
pile-up !

jet

asym !

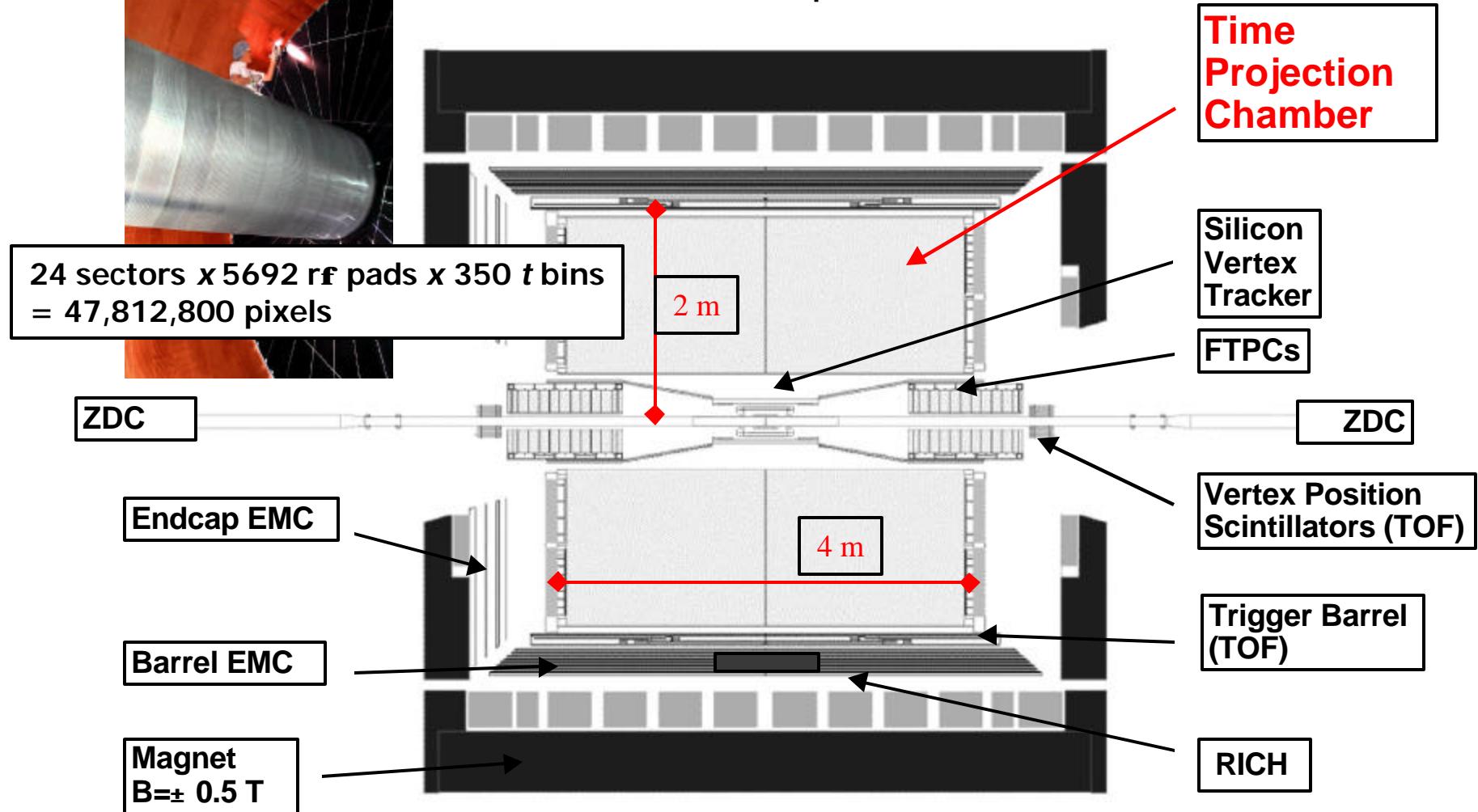
Are RHIC collisions dangerous ?

Not likely,
since the dawn of mankind we always had cosmic HI collisions.



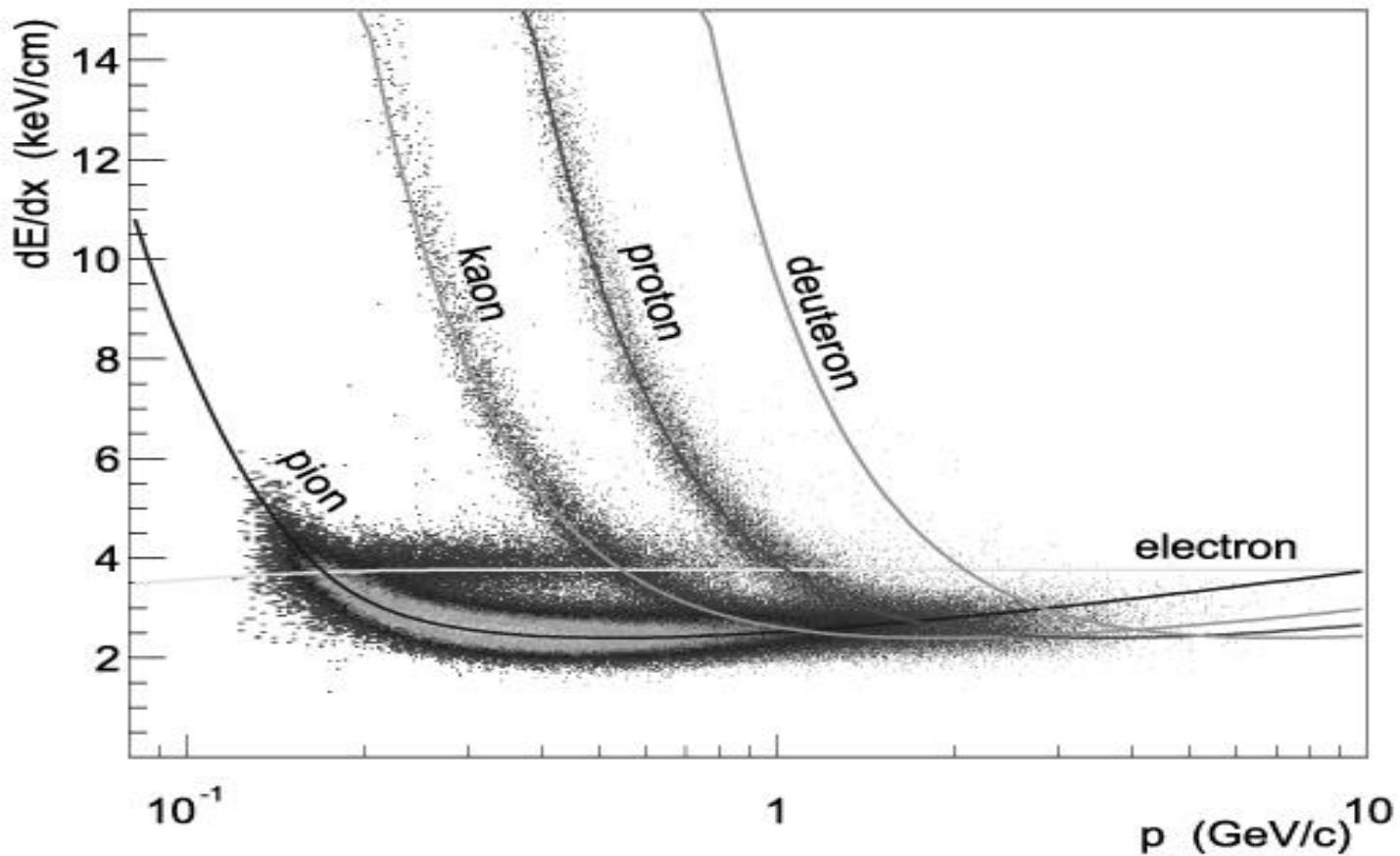


The STAR Experiment



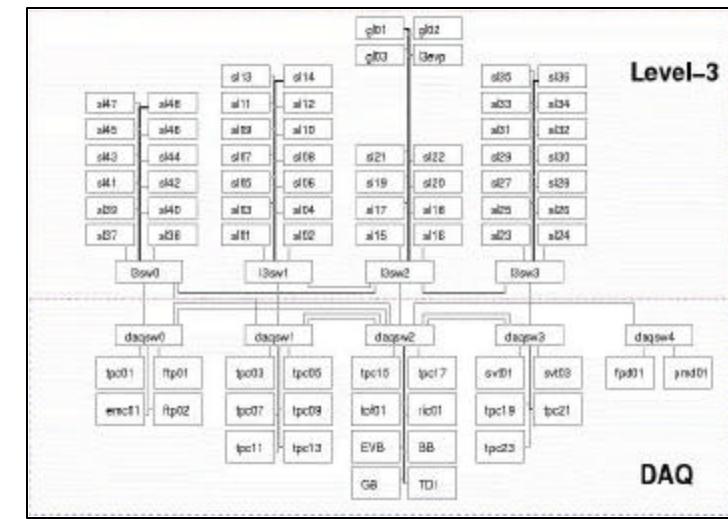
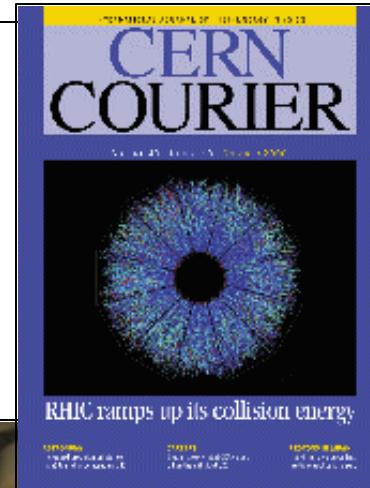
TPC dE/dx

- dE/dx resolution ~11% (offline ~8%)



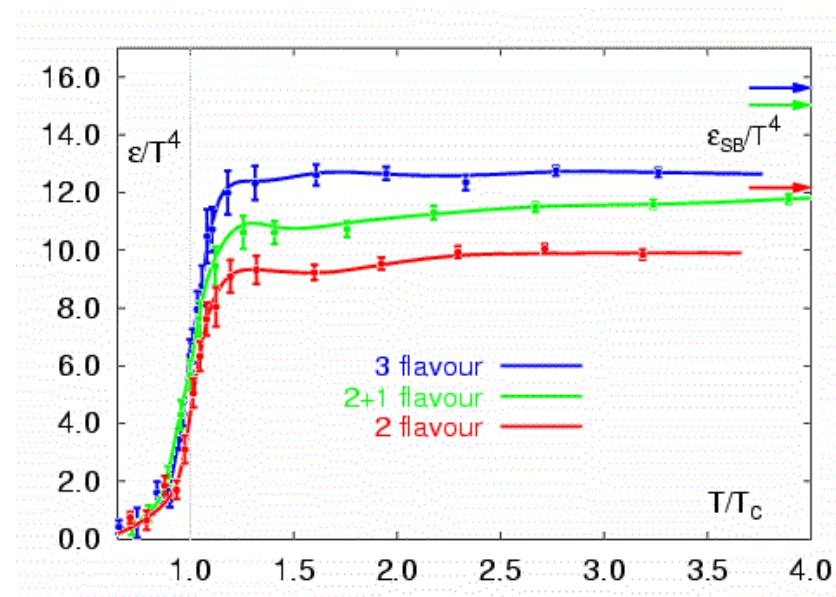
Level-3 Trigger System

- realtime full event reconstruction
(dE/dx , invariant mass)
- $t \leq 60$ ms
per 1 central Au+Au
 $N_{\text{Track}} = 4,500$
 $N_{\text{Cluster}} = 130,000$
- trigger rare events:
c,b quarks
(J/ψ , Upsilon)
Anti- ${}^3\text{He}$ / ${}^4\text{He}$

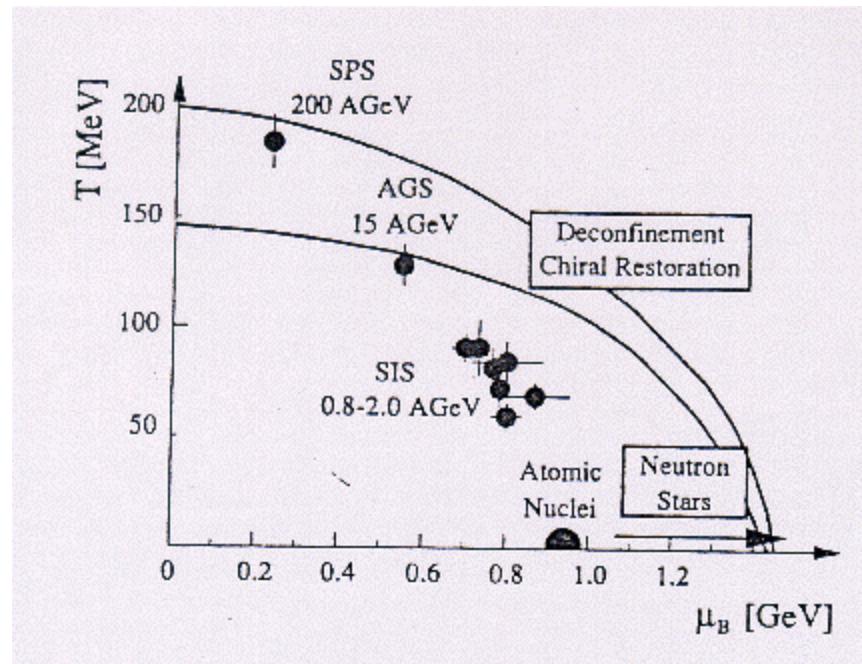


The QGP.

- deconfined (in QCD sense -> no bag anymore)
- 02/2000 CERN announced “QGP evidence”
 - strangeness enhancement (s-quark condensation in high T)
 - J/ψ suppression (gluon break-up)
- QCD phase transition
 - 1st order:**
 - mixed phase
 - hadrons with QGP bubbles
 - 2nd order:**
 - from pure 100% phase #1
 - to pure 100% phase #2
- predicted by lattice QCD
 $T_c = 160 \pm 3.5$ MeV



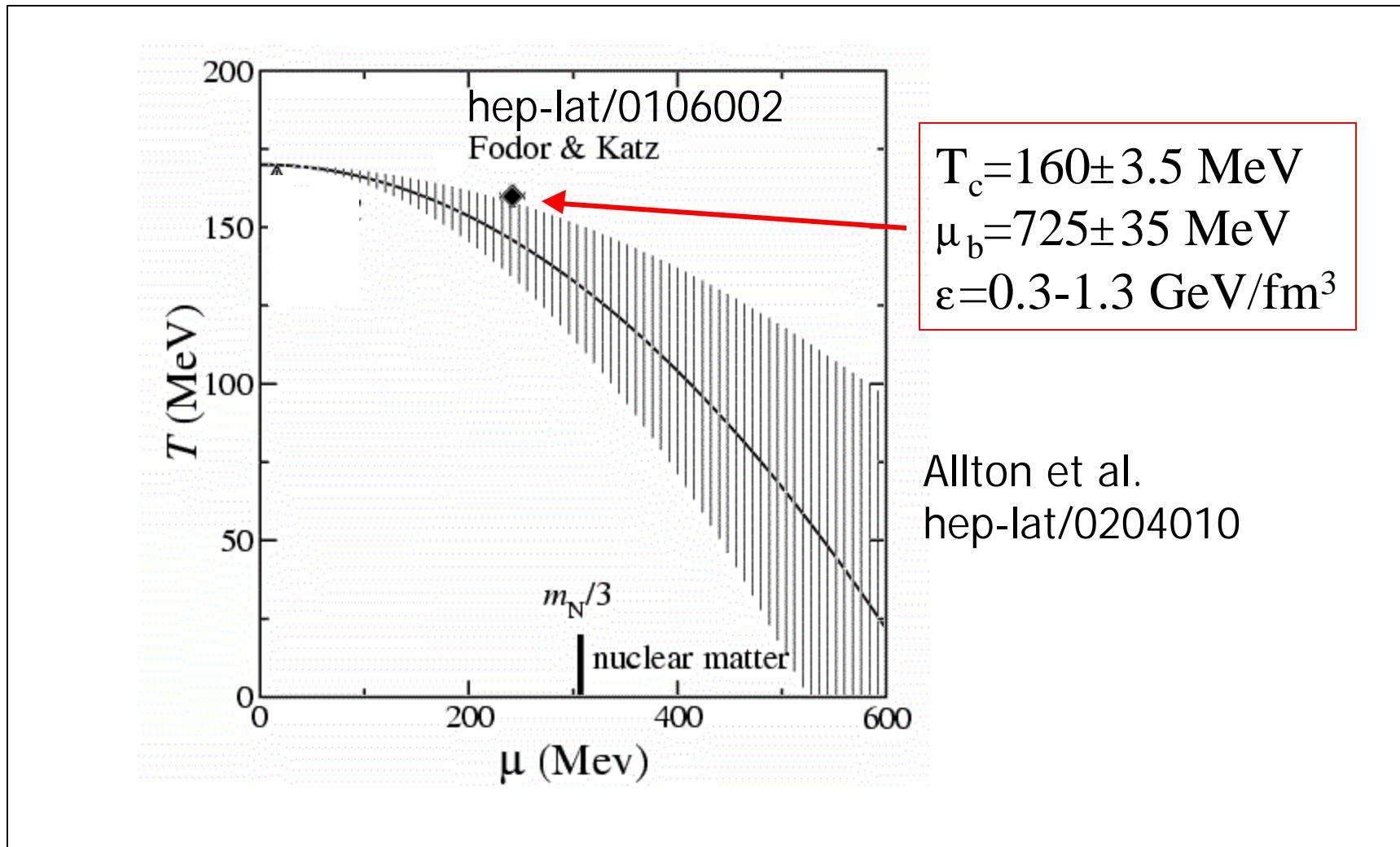
Phase Diagram before RHIC ?



$$\mu_B \sim N_{\text{Baryon}} - N_{\text{Anti-Baryon}}$$

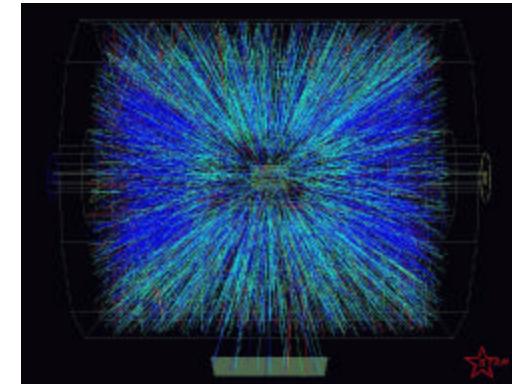
$$\mu_B \sim N_{\text{sea}} / N_{\text{valence}}$$

Where is the critical point ?



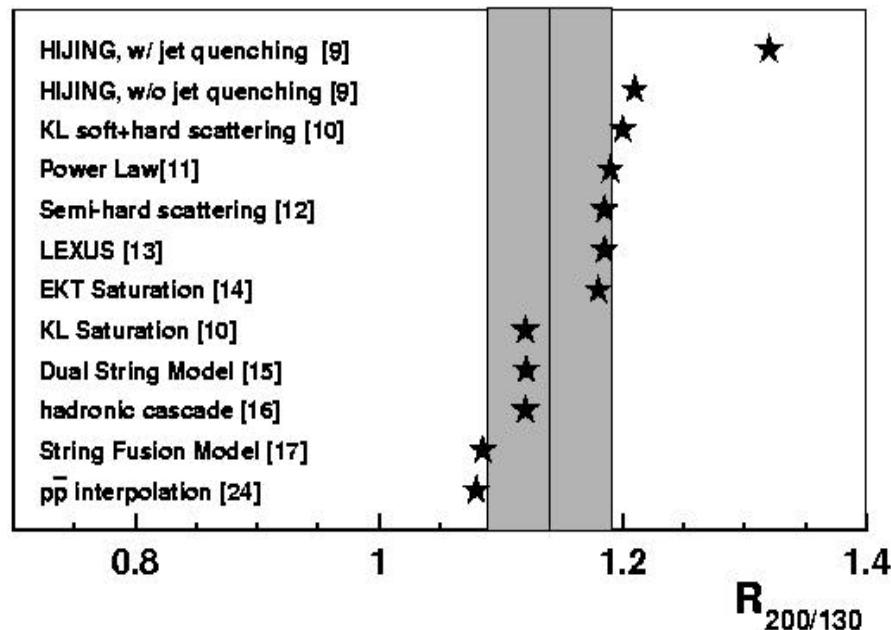
What did we learn from the number of particles ?

- ... 2-3 min after the 1st collision
- perturbative QCD expectation
 $N_{\text{charged}}(\sqrt{s}=200 \text{ GeV})$
= 1.14 $\times N_{\text{charged}}(\sqrt{s}=130 \text{ GeV})$
fits well -> no QGP needed for interpretation
- but pQCD models need either/or
 - "high density QCD"
 α_S^{eff} changes
 - "high temperature QCD"
 α_S^{eff} changes $\sim 1/(33-2N_f)\ln(T/T_{\text{crit}})$
both cases: gluons acquire "mass" (on-shell)
- no dramatic N_{charged} event-by-event fluctuations
 \rightarrow we are not near the critical point

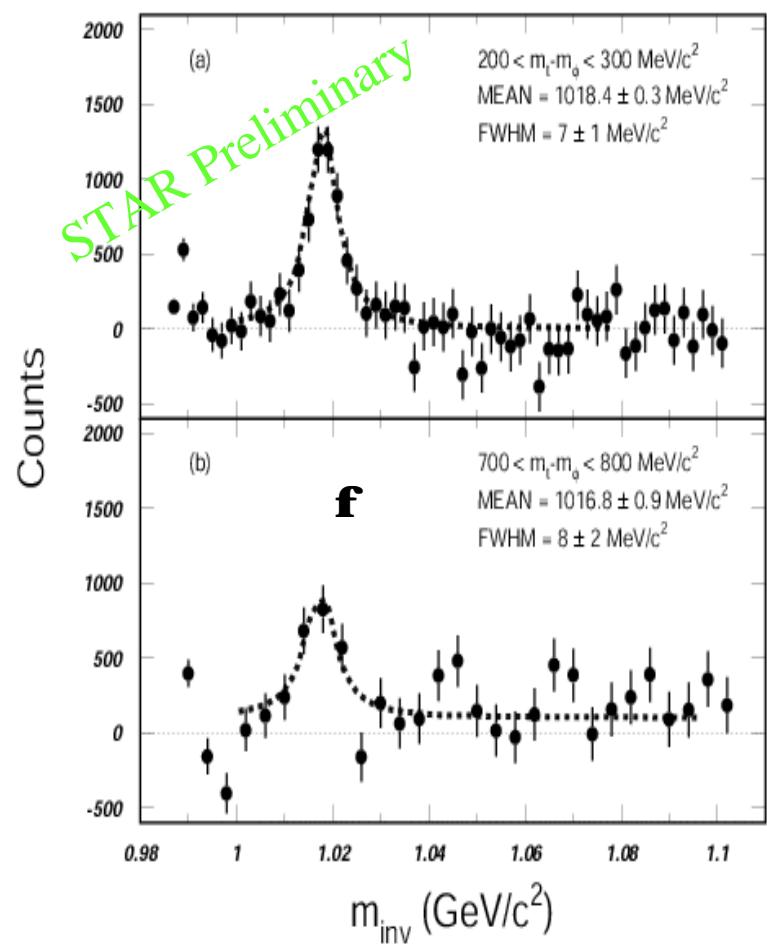
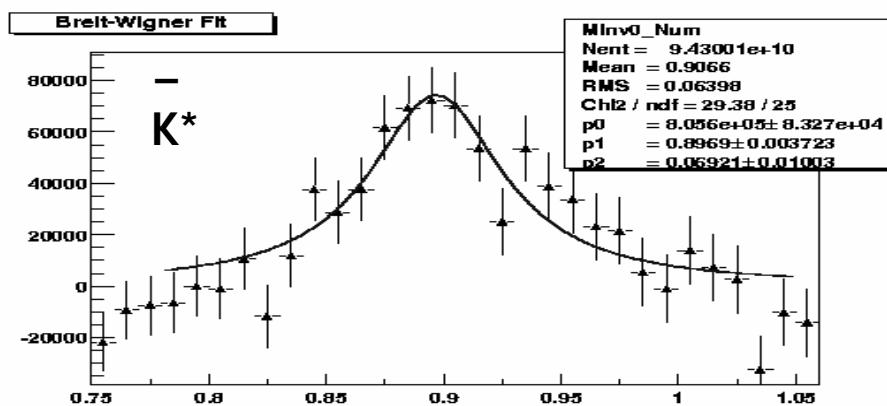
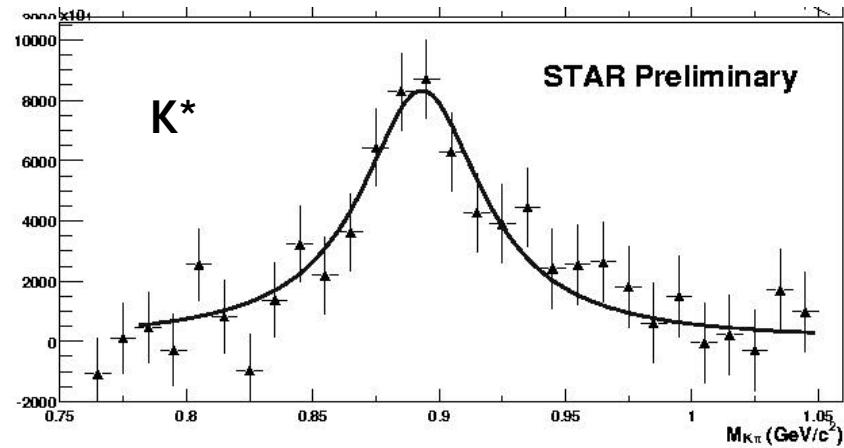


$$\frac{3\pi^2 \alpha_S A}{2Q^2_{\text{Saturation}}} \times \frac{x \cdot G(x, Q^2(x))}{\pi R_A^2} = 1$$

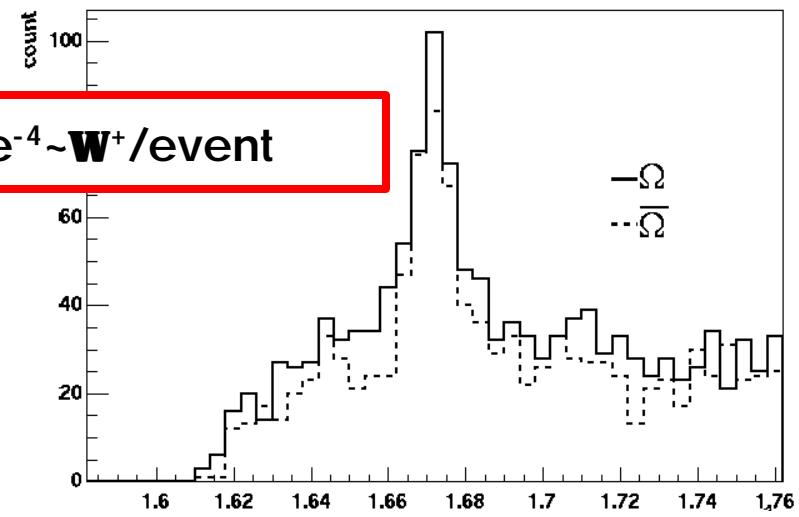
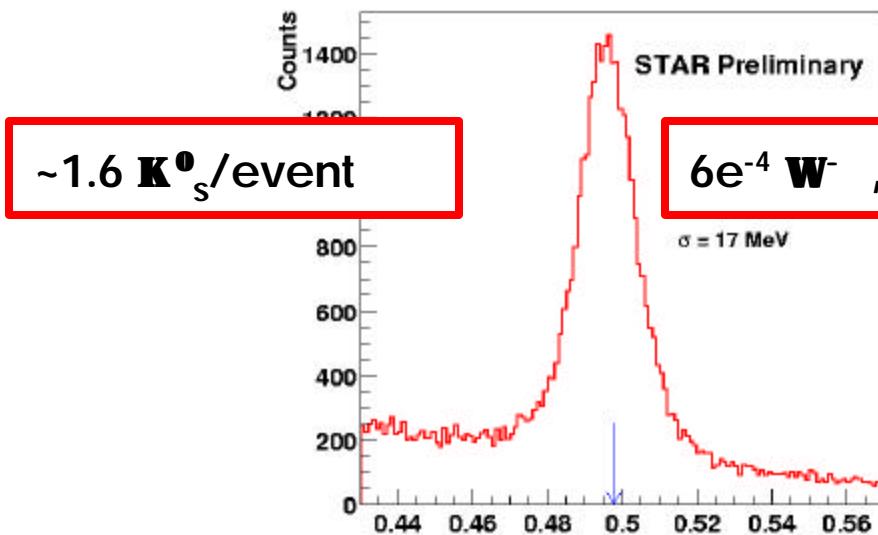
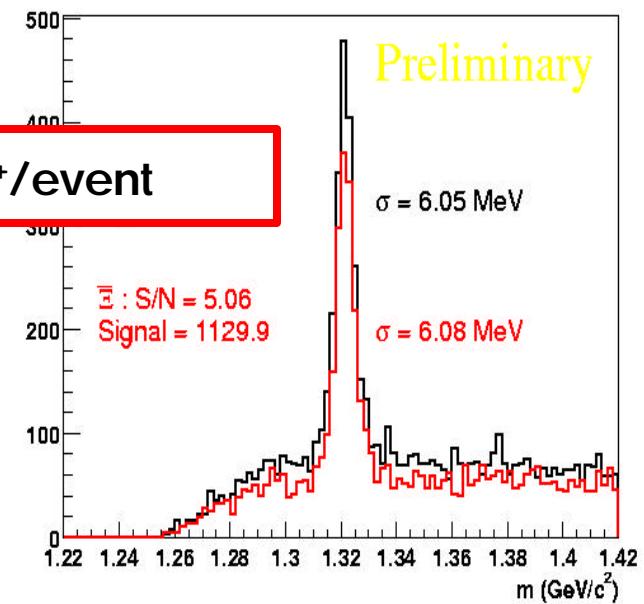
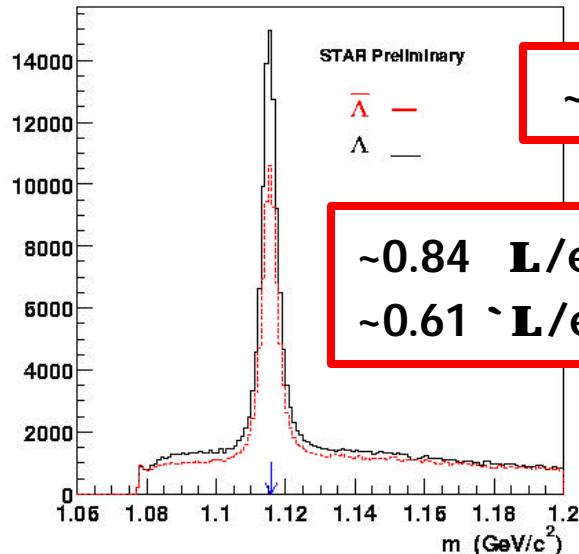
Which models fit, and which not ?



Data Quality 1: Mesons (qq)



Data Quality 2: Baryons (qqq)



Thermodynamics

- T from exponential fits to m_T

$$\frac{1}{m_T} \frac{dN}{dm_T} \propto A \exp\left(-\frac{m_T}{T}\right)$$

$$m_T = \sqrt{p_T^2 + m^2}$$

- μ_B from

$$\bar{p}/p \text{ ratio} = 0.65 \pm 0.03(\text{stat}) \pm 0.03(\text{syst}) = \lambda_q^{-6}$$

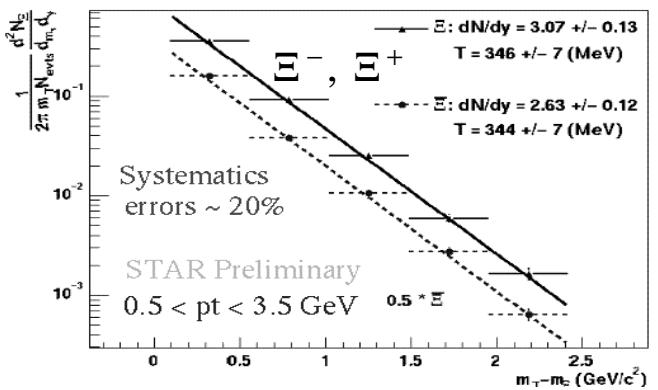
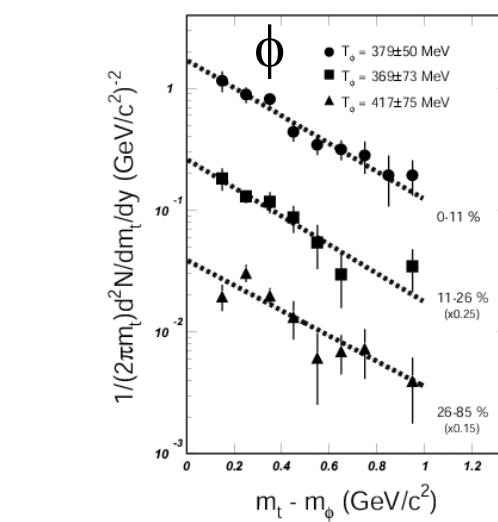
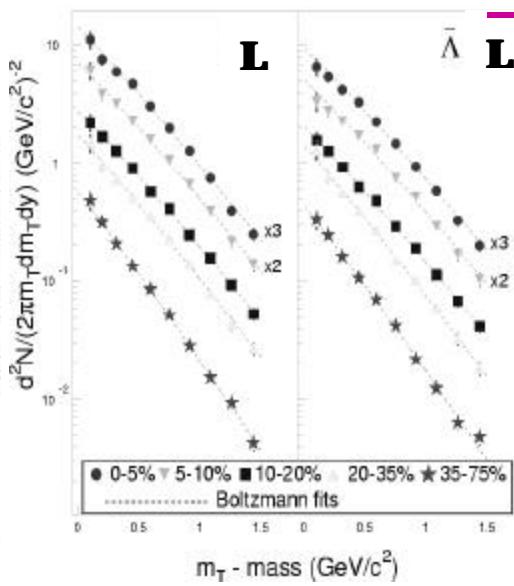
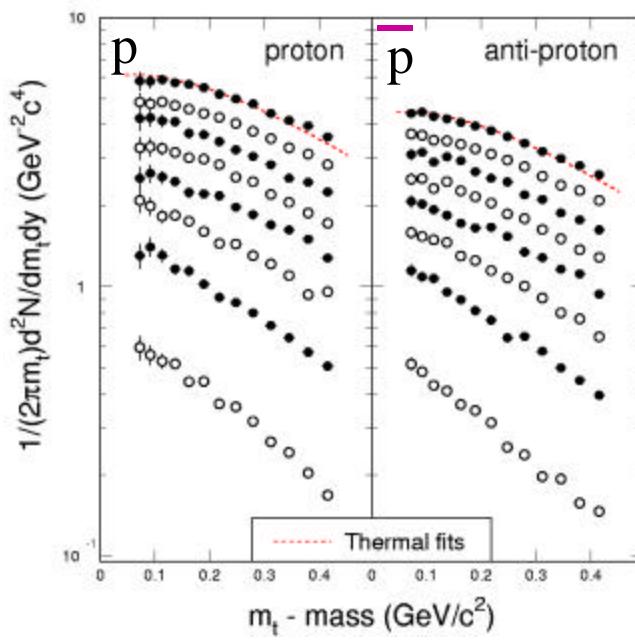
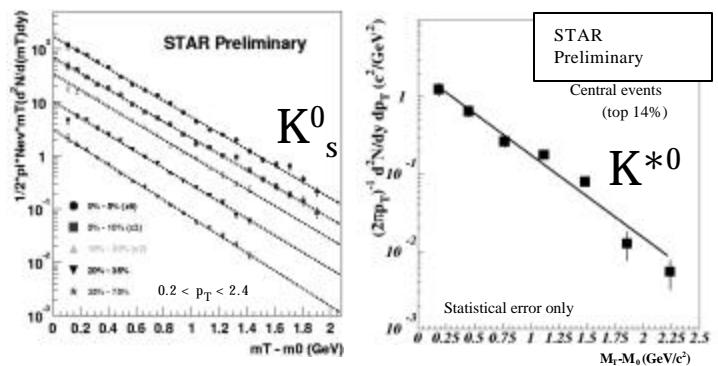
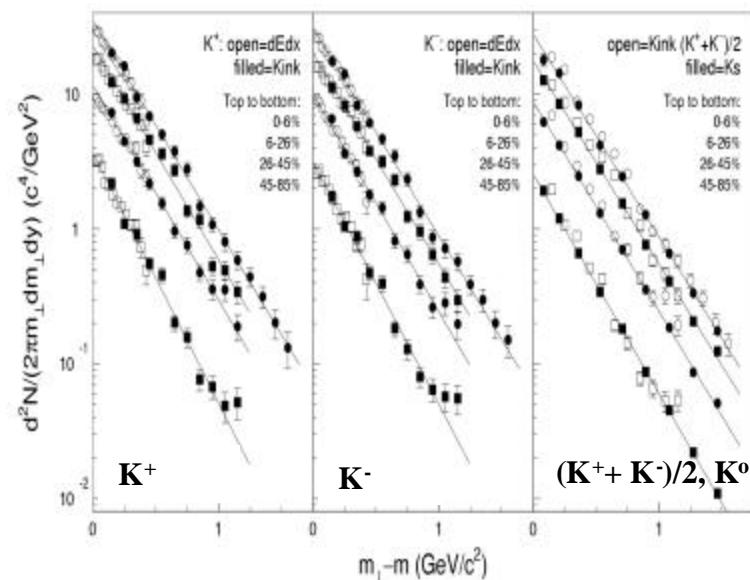
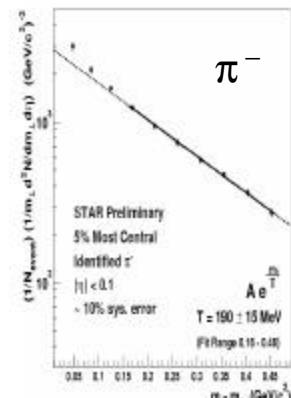
$$\lambda_q = 1.074$$

$$\lambda_q = \exp(\mu_q/T)$$

$$\mu_b = \mu_u + 2\mu_d = 40.1 \text{ MeV}$$



m_T Distributions



Thermal fit Result:

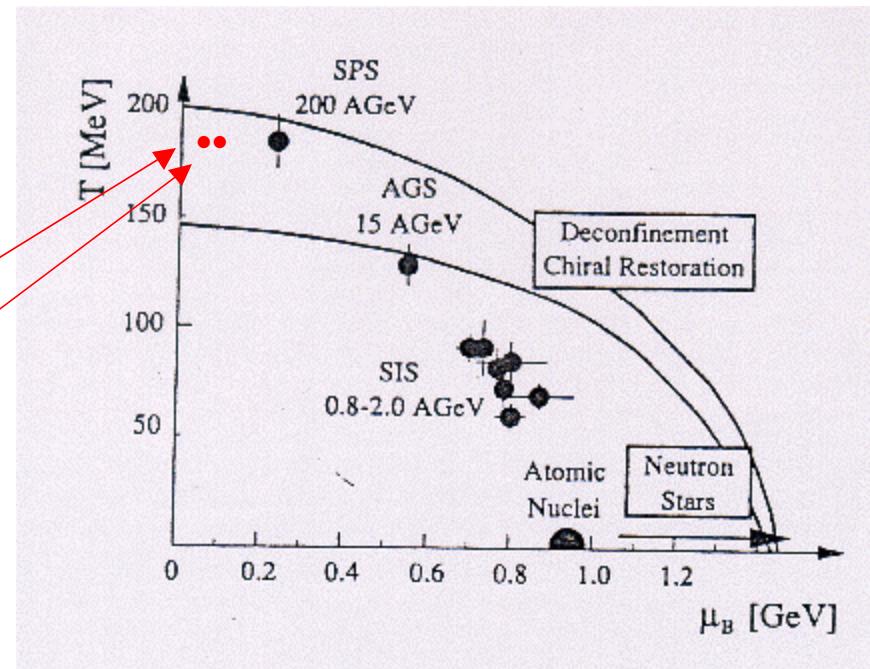
$T = 176 \text{ MeV}$ (130 GeV), $T = 177 \text{ MeV}$ (200 GeV)

$$\mu_B = 41 \text{ MeV} \text{ (130 GeV)}$$

$$\mu_B = 29 \text{ MeV} \text{ (200 GeV)}$$

$$T = 2.1 \cdot 10^{12} \text{ K}$$

200 GeV
130 GeV



Sun
 $15.6 \cdot 10^6 \text{ K}$
Supernova
 $\sim 10^9 \text{ K}$
Plasma fusion
 $55 \cdot 10^6 \text{ K}$
Laser fusion
 $4 \cdot 10^6 \text{ K}$

Universe:

$$T_{\text{Planck}} = 1.4 \cdot 10^{34} \text{ K}$$

but maybe Hagedorn-limited $\sim 1/R$

$$\mu_B \sim N_{\text{Baryon}} - N_{\text{Anti-Baryon}}$$

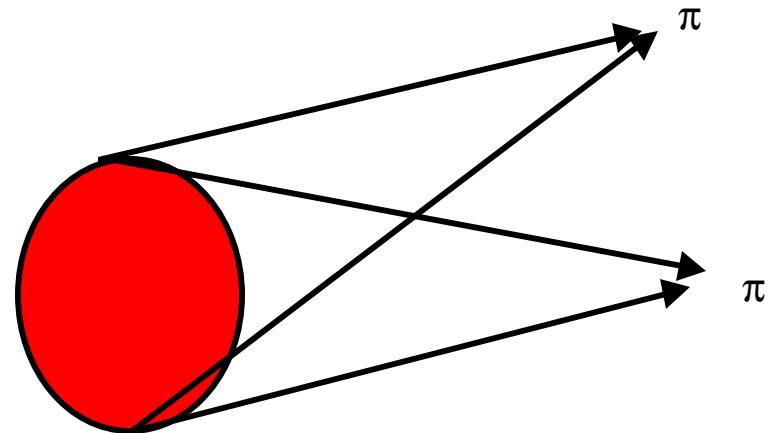
$$\mu_B \sim N_{\text{sea}} / N_{\text{valence}}$$

What is the size ?

- Hanbury-Brown-Twiss (HBT) Interferometry
 $\pi\pi$ pairs (bosons) close in Δp

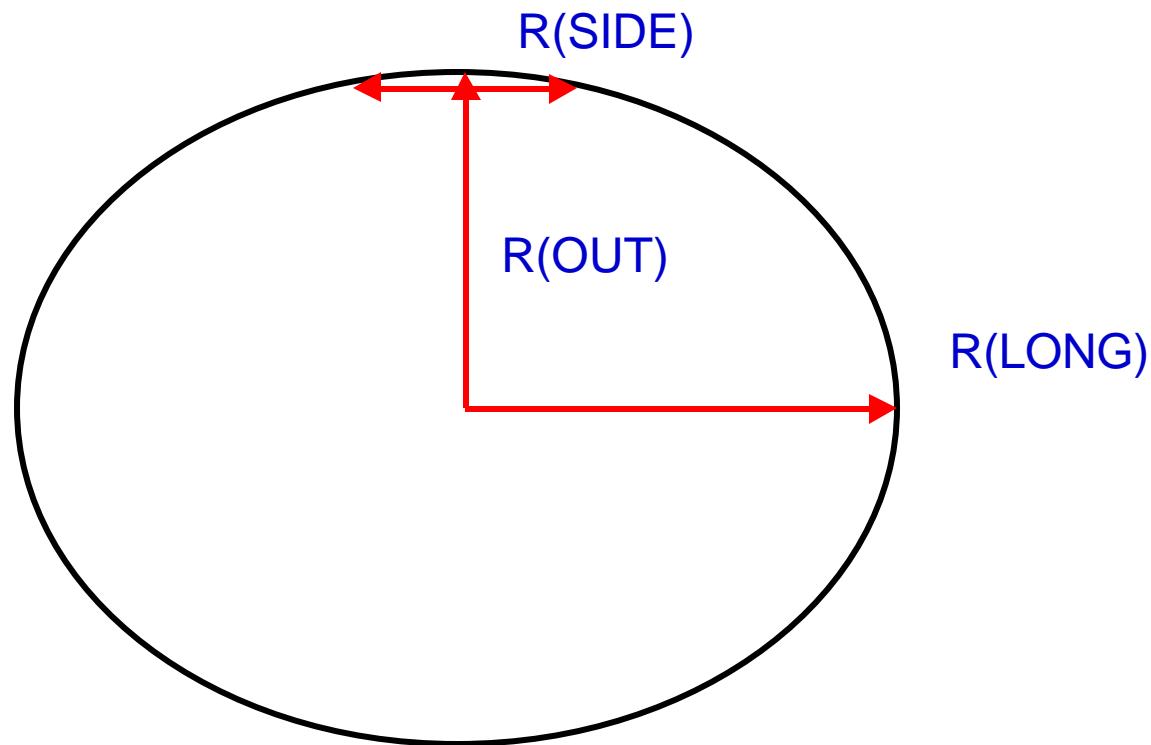
- Universe: $R(\text{galaxy})$
difference in arrival time

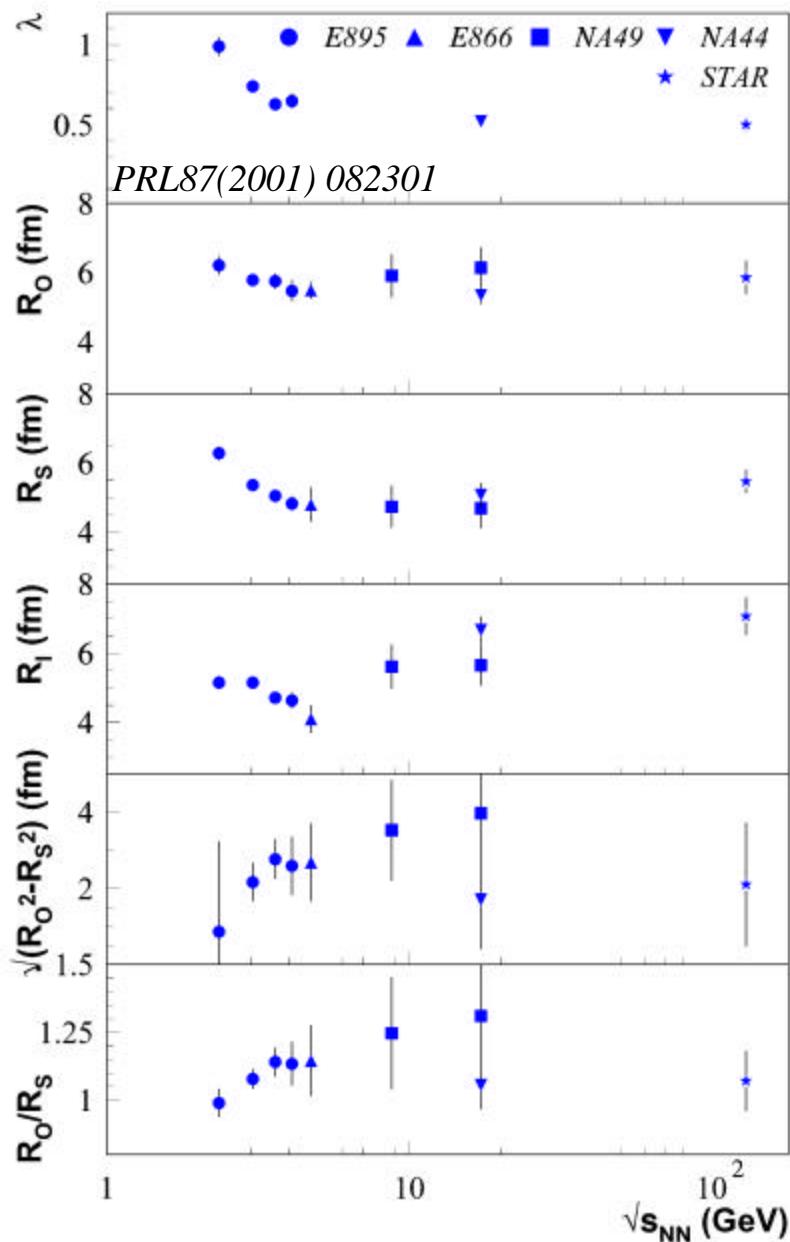
- AA collisions:
difference in momentum Δp
Fourier Transform $\Delta p \rightarrow \Delta x$
Correlation Function
 $C(\Delta p) = 1 + \lambda \times \text{FourierTransform}[\rho(\Delta p)] = 1 + \lambda \exp(-\Delta p^2 \Delta R^2)$
(λ for resonances)



- remember: expanding system $R=R(t)$!
- Lorentz-boost: sphere \rightarrow cylinder

Expanding Volume



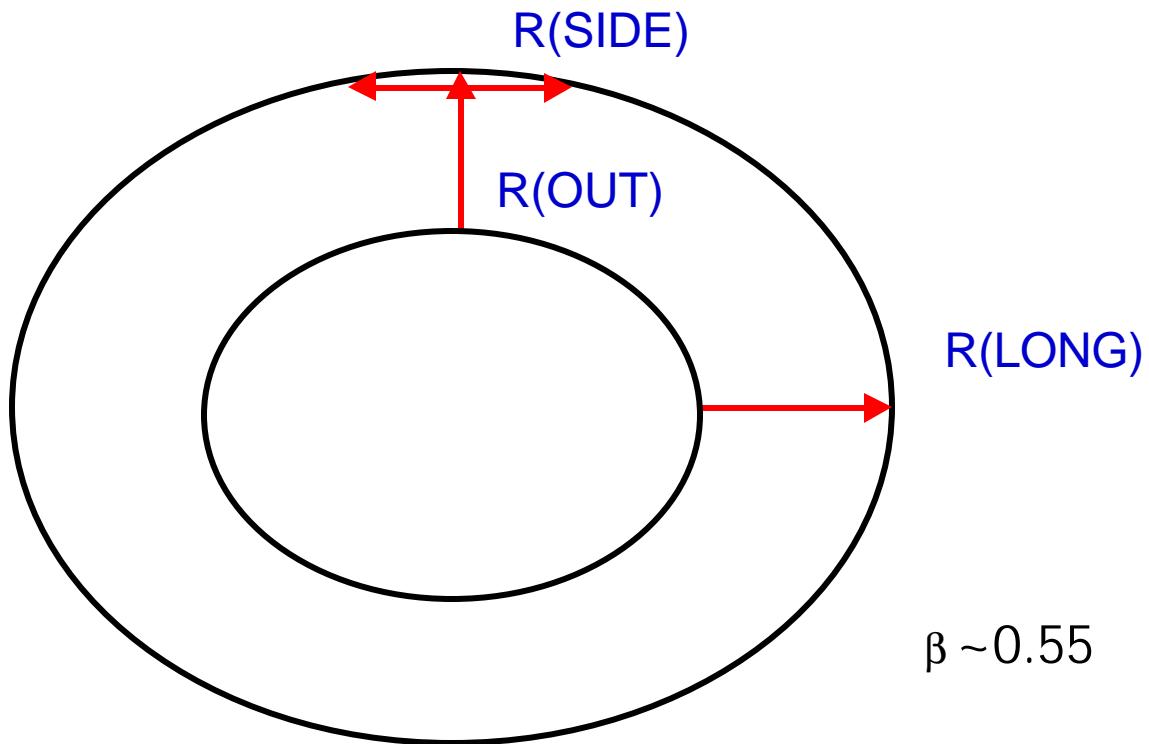


Pion HBT vs. σ_s

World compilation + STAR

- **Surprising:** Size roughly same AGS = SPS = RHIC $R < 10$ fm
- changes with beam energy negligible < 0.5 fm
- R increase with centrality (\sim overlap region) \rightarrow O.K.
- **Unexpected:**
 - short “freeze-out”
 - explosive source

$R_{\text{out}}/R_{\text{side}} \gg 1$: Expanding Shell ("blast wave")



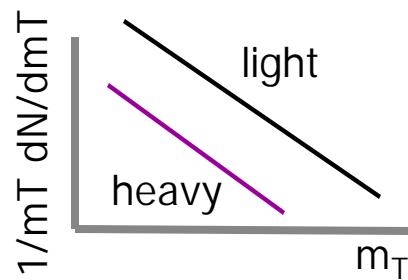
Universe:

at $T \sim 100$ MeV horizon distance $L \sim 10$ km !

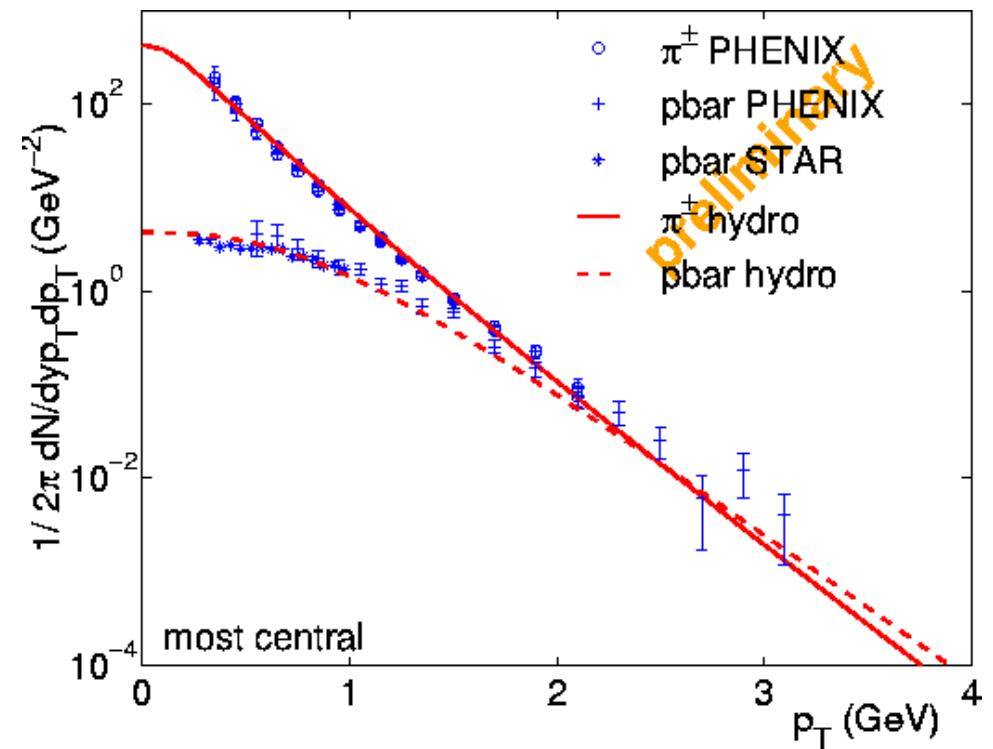
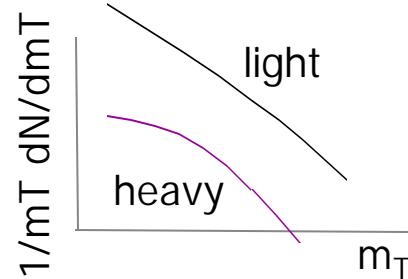
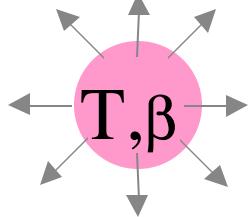
expansion rate slower $\sim 10^{19}$ ($= 1/M_{\text{Planck}}^2$)

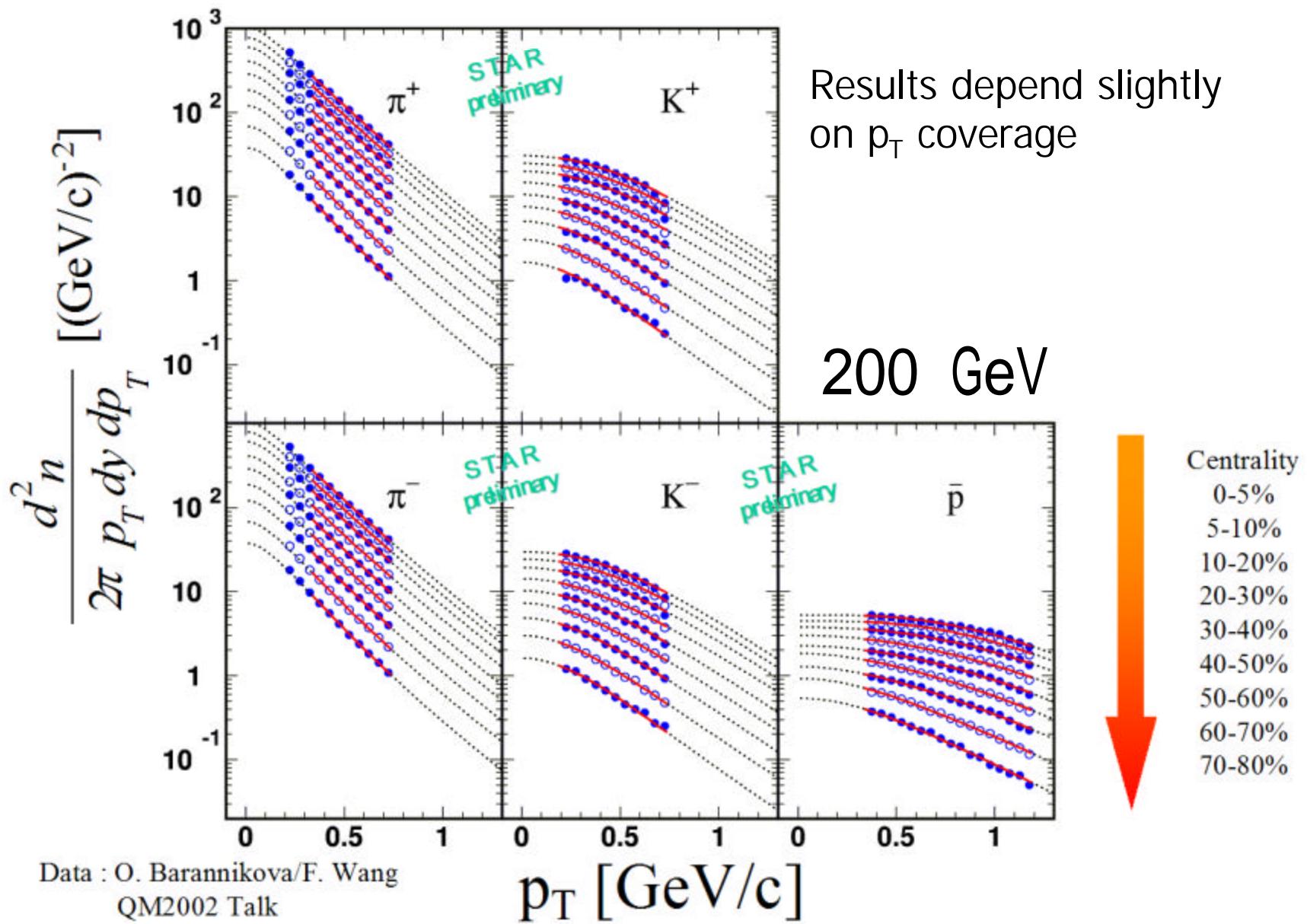
Is it an explosion ?

purely thermal source



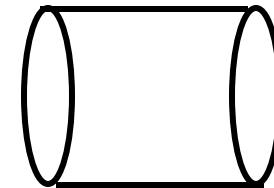
explosive source





What is the quark/gluon density @ t=0 ?

- we got N_{charged}
- we got R
assume cylinder, lorentz-invariant rapidity y
 $t=0 \rho=\infty$, but @ $t=0.2 \text{ fm}/c$:
 $\rightarrow \rho = 20/\text{fm}^3 = 15 \times \rho[\text{cold Au}]$
(hadrons definitely in-existent)

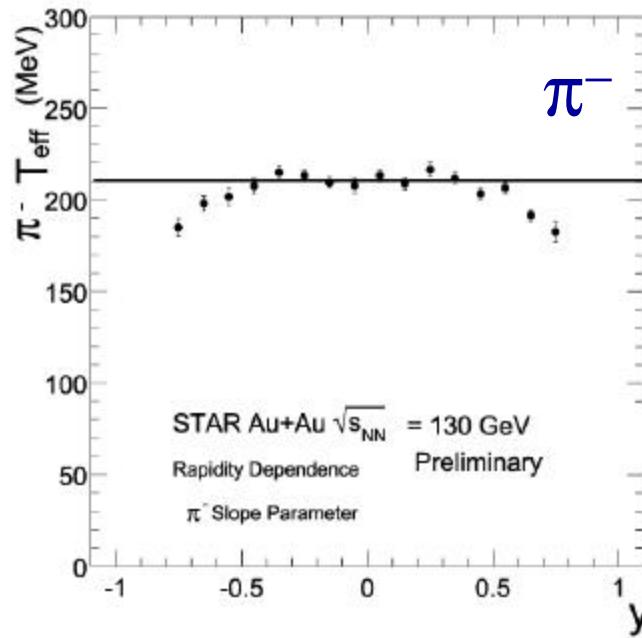


$$\rho \approx \frac{dN}{dy} \frac{1}{\pi R^2 t}$$

- What is the energy density ?
we got average momentum $\langle p \rangle$
90% π , so assume m_π , then $E^2=m^2+p^2$
 $\rightarrow \varepsilon \sim 5 \text{ GeV/fm}^3 = 30 \times \varepsilon[\text{cold Au}]$

better pQCD estimate: $\varepsilon \sim 18 \text{ GeV/fm}^3$ Phys. Lett. B 507(01)121
SPS $\varepsilon \sim 3 \text{ GeV/fm}^3$, predictions before RHIC start up to 30 GeV/fm³

Boost-Invariance ?



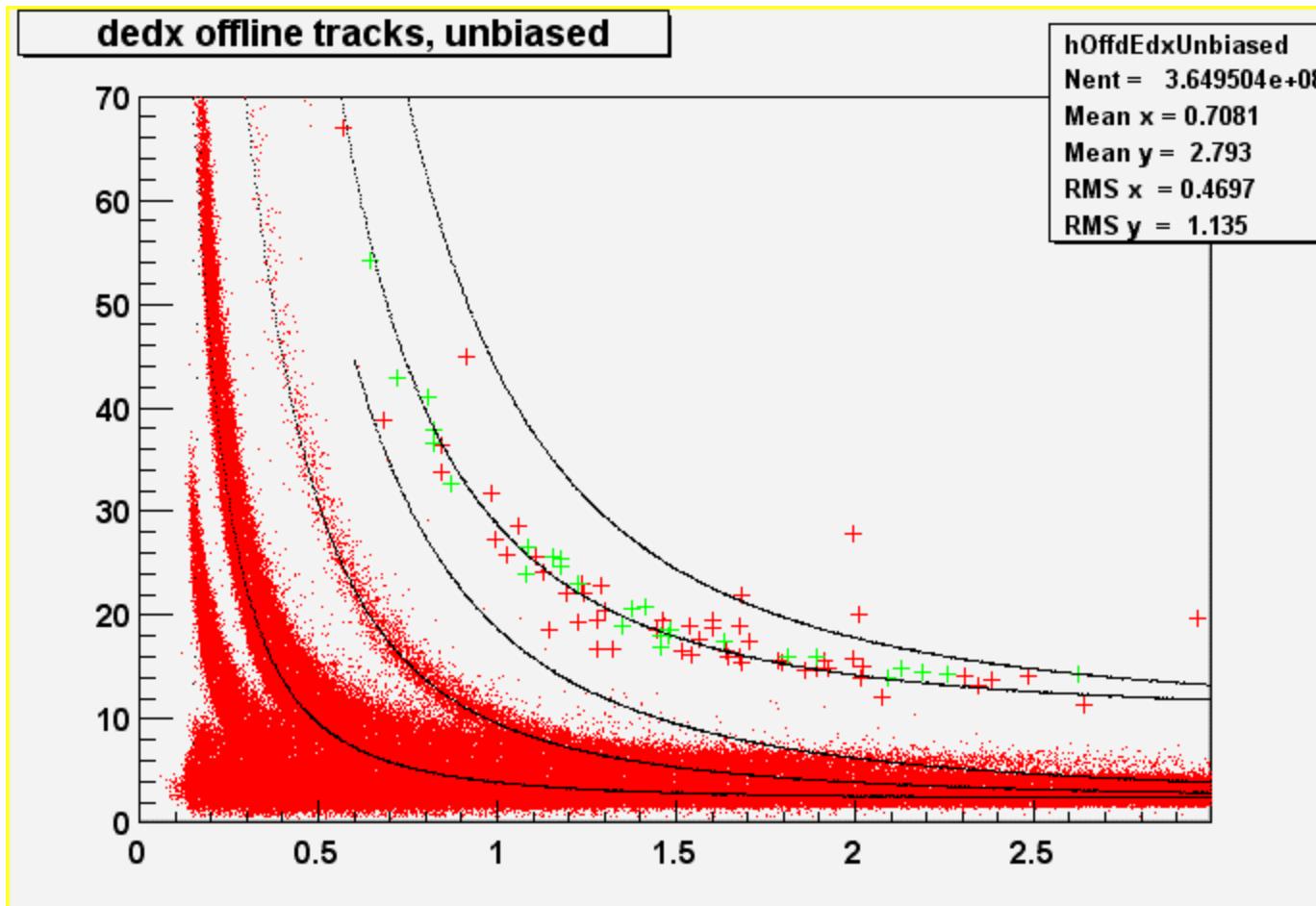
Boost invariance only achieved in small region $|y|<0.5$

- No surprise for protons:
p/p \sim 0.65 means 2/3 come from pair production,
but 1/3 comes from Au nuclei (de-accelerated)
- But surprise for pions !

How do we create Anti-He ?

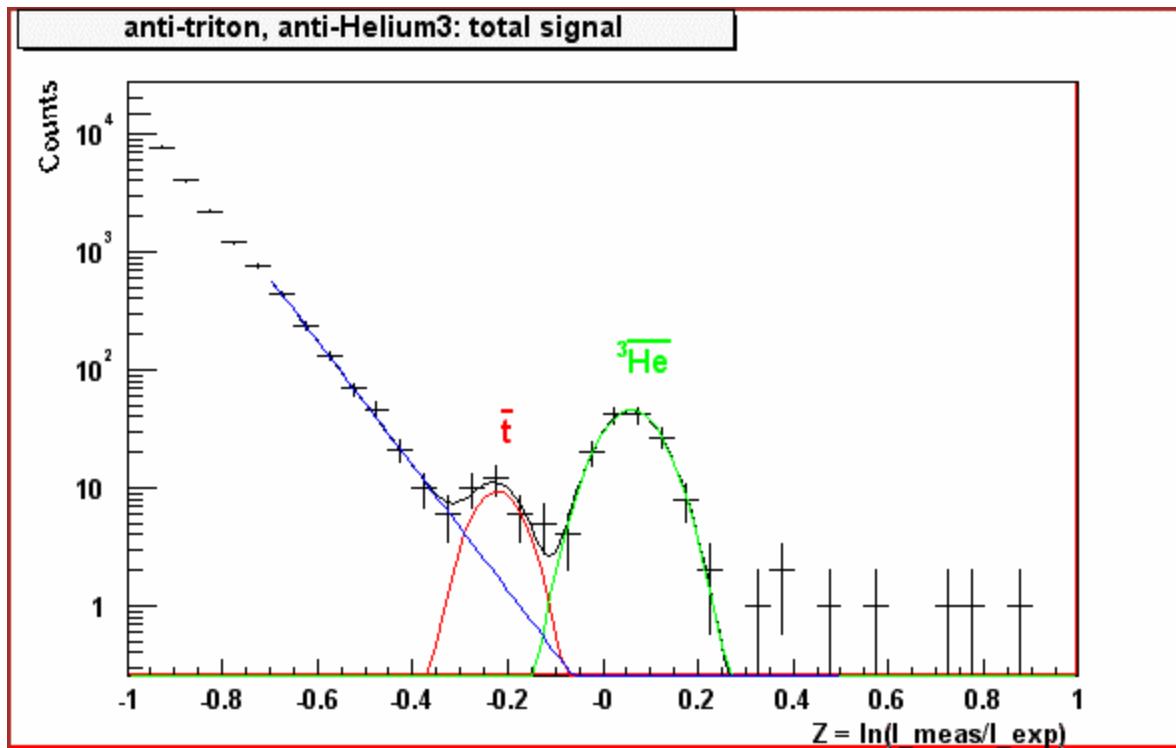
- cluster of particles can also be beam remnants,
but cluster of anti-particles cannot
- Coalescence:
N protons + overlapping wave functions
(Δx and Δp small)
- Ansatz $N_{\text{deuteron}} = f \times N_{\text{proton}}$
“penalty factor” $f \sim 1/2500$
- Level-3 trigger system, ~4 months:
charge = -2
(almost background free, but rare) needs $15 \cdot 10^6$ events
- total sample 2002
 - 160 Anti-He³
 - no Anti-He⁴ yet (would be first observation)

Level-3 Q=-2 Trigger

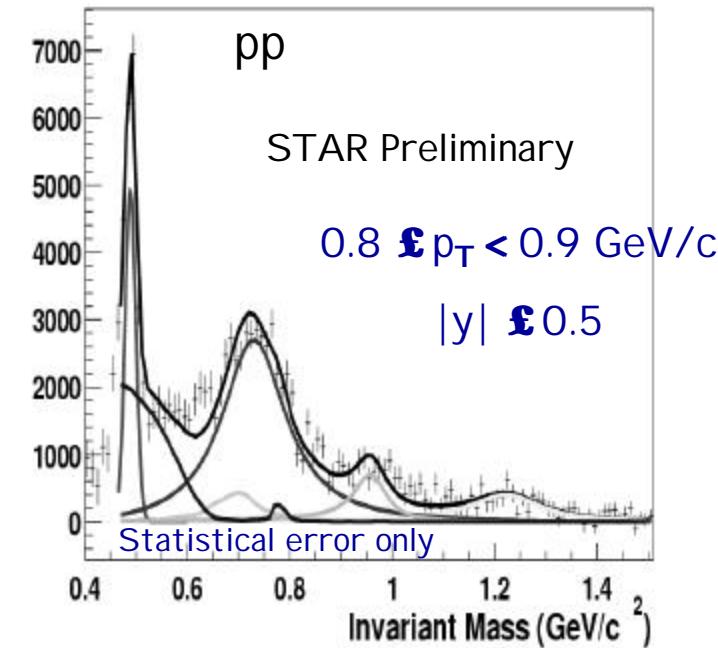
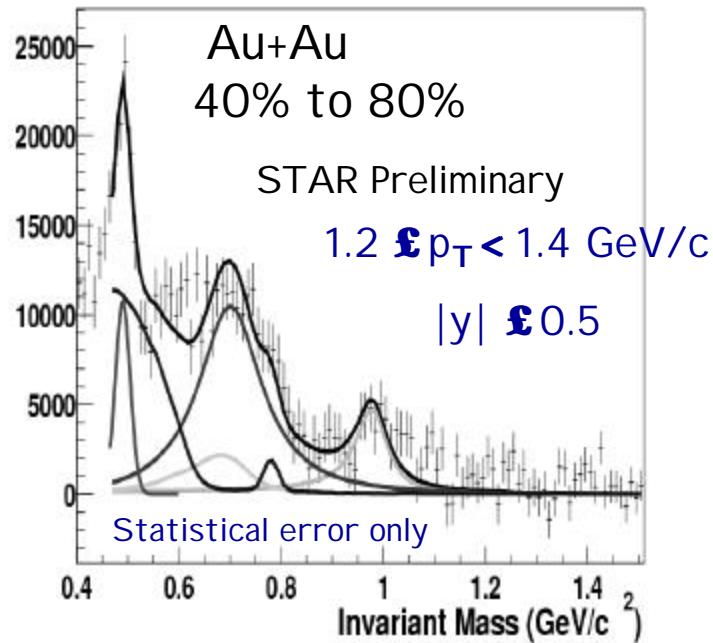


Q=-2 Candidates

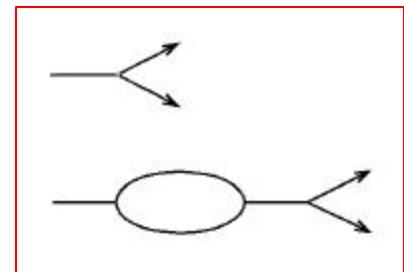
- Result: ψ overlap volume: $R^3 = 107 \pm 7 \text{ fm}^3 \rightarrow R = 4.8 \text{ fm}$
but proton/neutron freeze-out late $t \sim 15-20 \text{ fm/c}$
 \rightarrow early coalescence (parton vs. baryon coalescence ?)



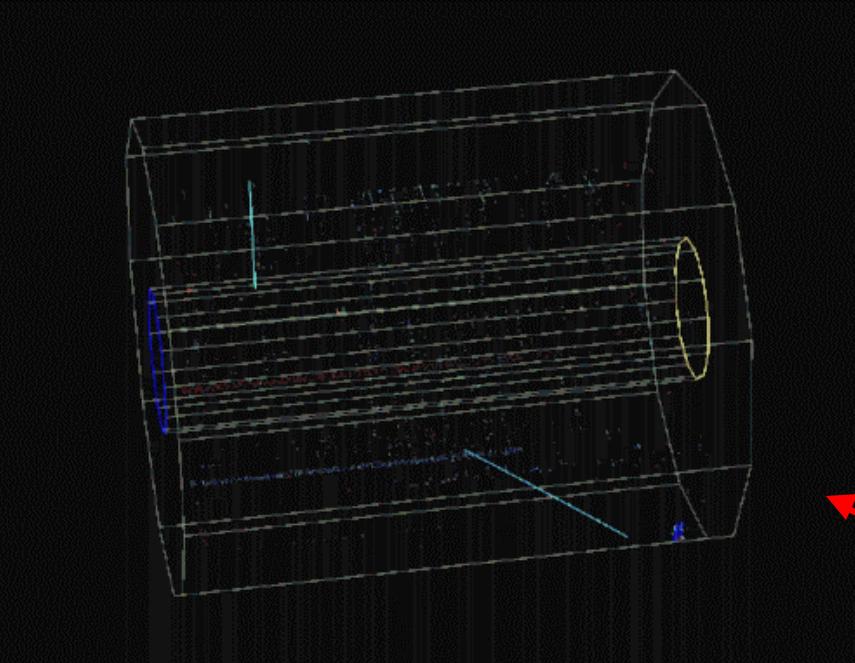
Next step after $T(\pi)$ and Radius($\pi\pi$) ? $\pi\pi$ resonances.



- ρ decay ($\Delta t = 1.3 \text{ fm}$) < π freeze-out ($\Delta t \sim 10 \text{ fm}/c$) < f_0 decay ($\Delta t \sim 15 \text{ fm}$)
- ρ mass appears shifted 20-80 MeV = $f(p_T)$
- + phase-space peak shape ("re-scattering volume")
- + re-scattering (QM interference, $d \sim 1/3 \text{ fm}$) seems to shift mass back to PDG value



Ultra-peripheral Au+Au (nuclei don't touch)



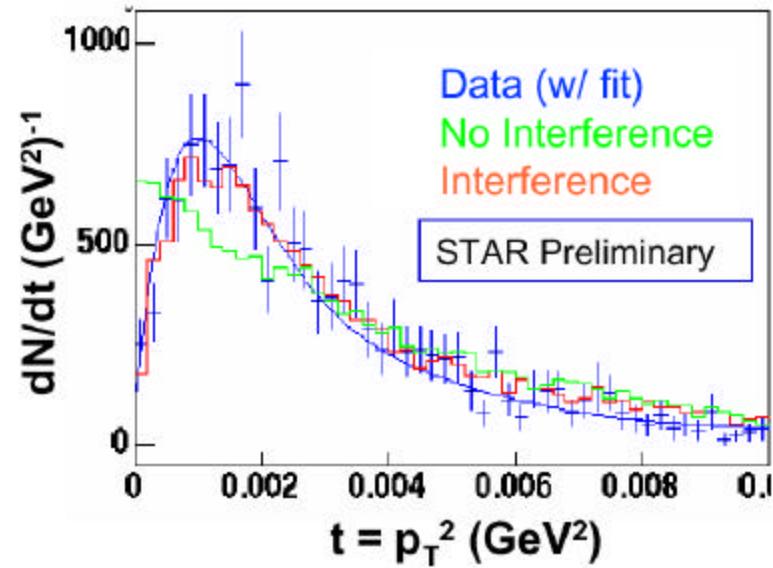
$b > R(\text{Au})$
(Weizsäcker-Williams,
virtual quanta)
 $\text{Au} + \text{Au} \rightarrow \text{Au}^* + \text{Au}^* + \rho^0$

γ pomeron $\rightarrow \rho^0 \rightarrow \pi^+ \pi^-$
1st event 08.24.2000



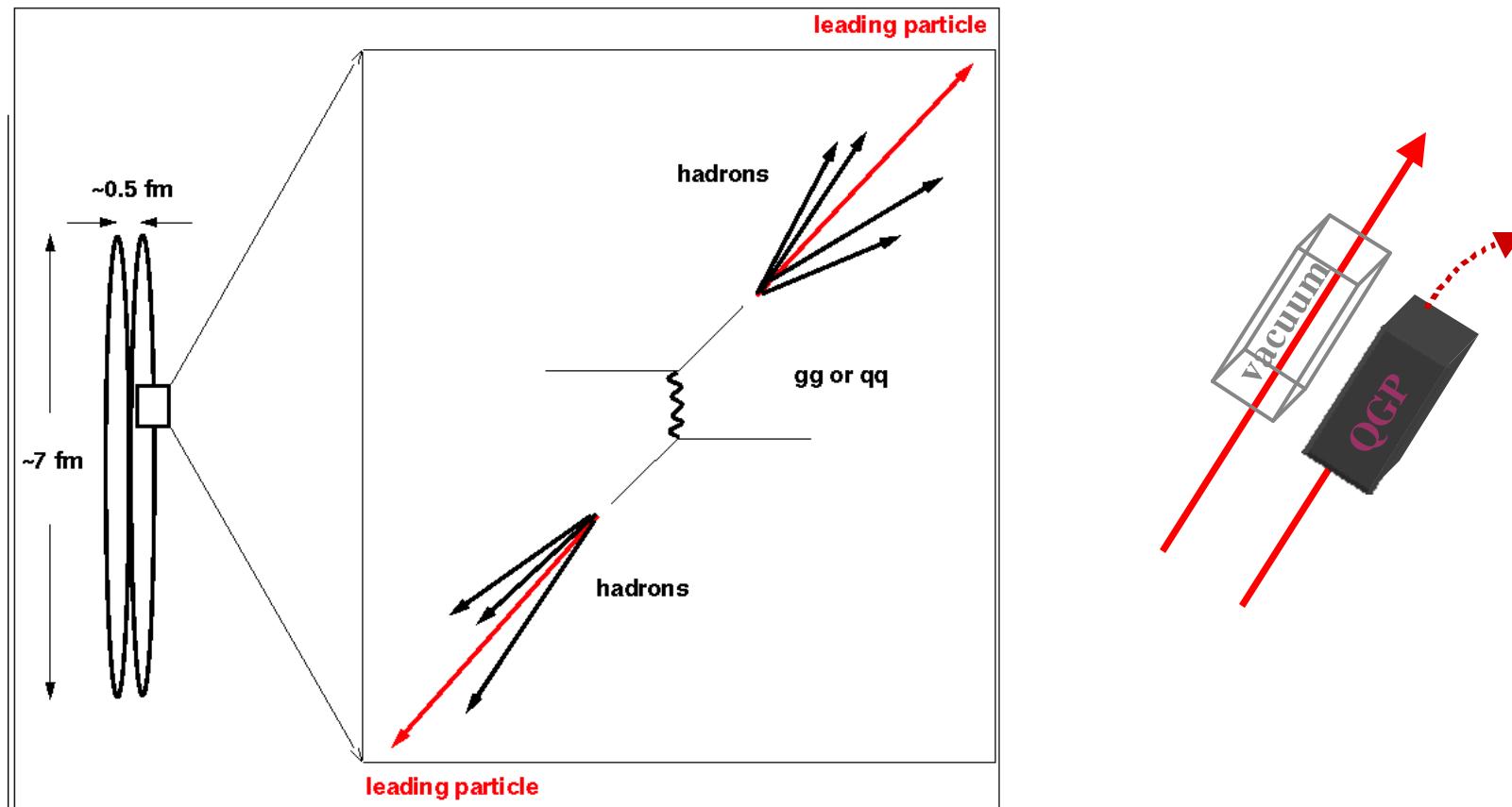
QM non-local entanglement in Au^{*}Au^{*}, $\mathbf{r}^o \circledR \mathbf{p}^+ \mathbf{p}^-$

- indistiguishable initial states (Au^{*} Au vs. Au Au^{*})
 $\Psi_{\pi\pi}$ contains destructive interference (parity -)
 $\sigma = \sigma_0(1 - \cos(\mathbf{p} \cdot \mathbf{b}))$
- $p_T < < 1/\langle b \rangle$ with $\langle b \rangle \sim 20$ fm
 $c\tau (\rho^o) = 1.32$ fm/c
→ $\pi\pi$ wave function non-local !
- coherence preserved even after the ρ decay
- clear “Einstein-Podolsky-Rosen” phenomenon



Jets in AA

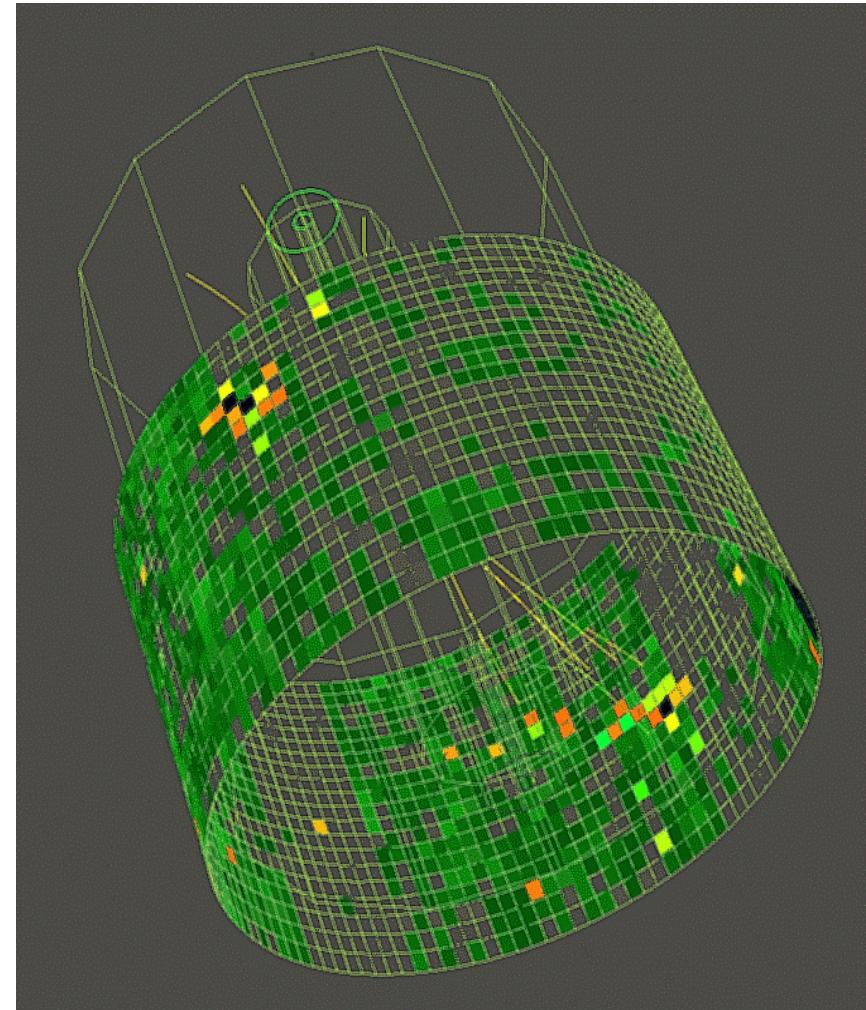
- $\sqrt{s}(\text{RHIC}) \sim 10 \cdot \sqrt{s}(\text{SPS})$ - for the first time σ_{jet} accessible in AA collisions
- gg or qq scattering is very early $t \sim 0.1 \text{ fm}/c$ (cylinder very short)
- pQCD $\sqrt{s}=200 \text{ GeV}$: ~ 405 gluons, ~ 132 quarks, ~ 38 anti-quarks



Jets at STAR (p_T , η =200 GeV)

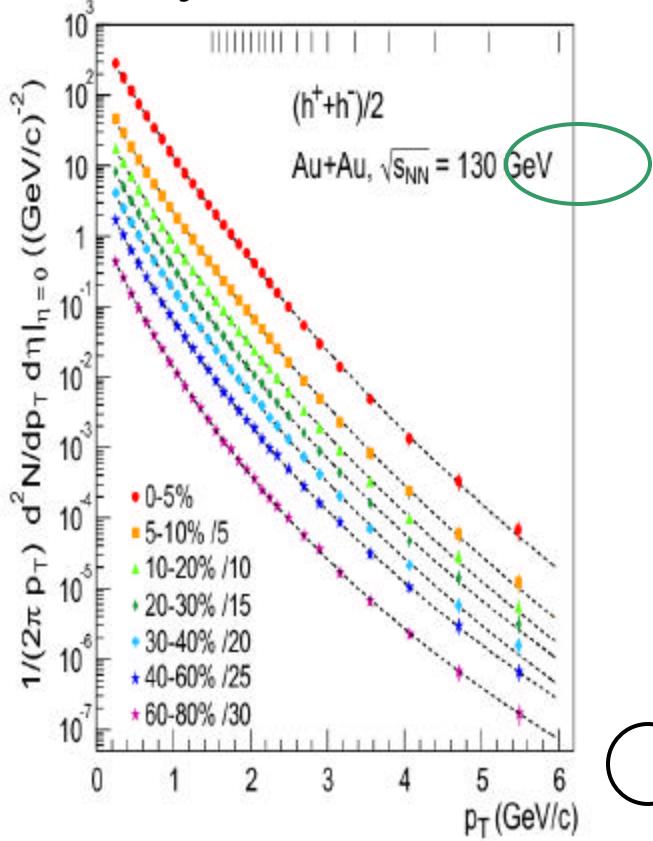
2 jet (gg or qq)
all tracks $p_T > 1$ GeV/c

What about Au+Au ?
pQCD estimate
 $E_T > 1$ GeV
 $N_{\text{jet}} \sim 500$
almost impossible
for jet finder algorithm.

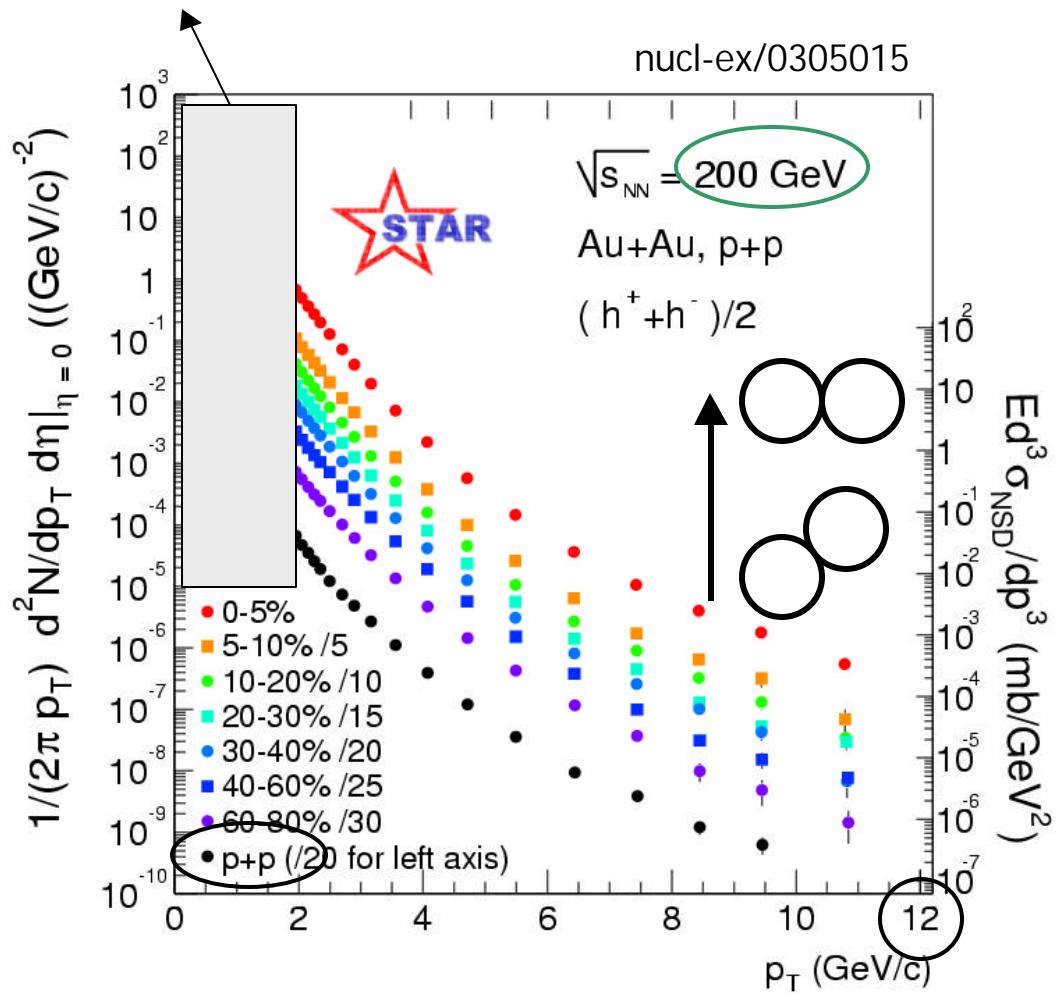


Charged high pT hadrons = leading jet particles

Phys. Rev. Lett. 89, 202301



99.5%



pp measured in same detector.

Jens Sören Lange (Frankfurt)

Search for New QCD Phenomena at STAR

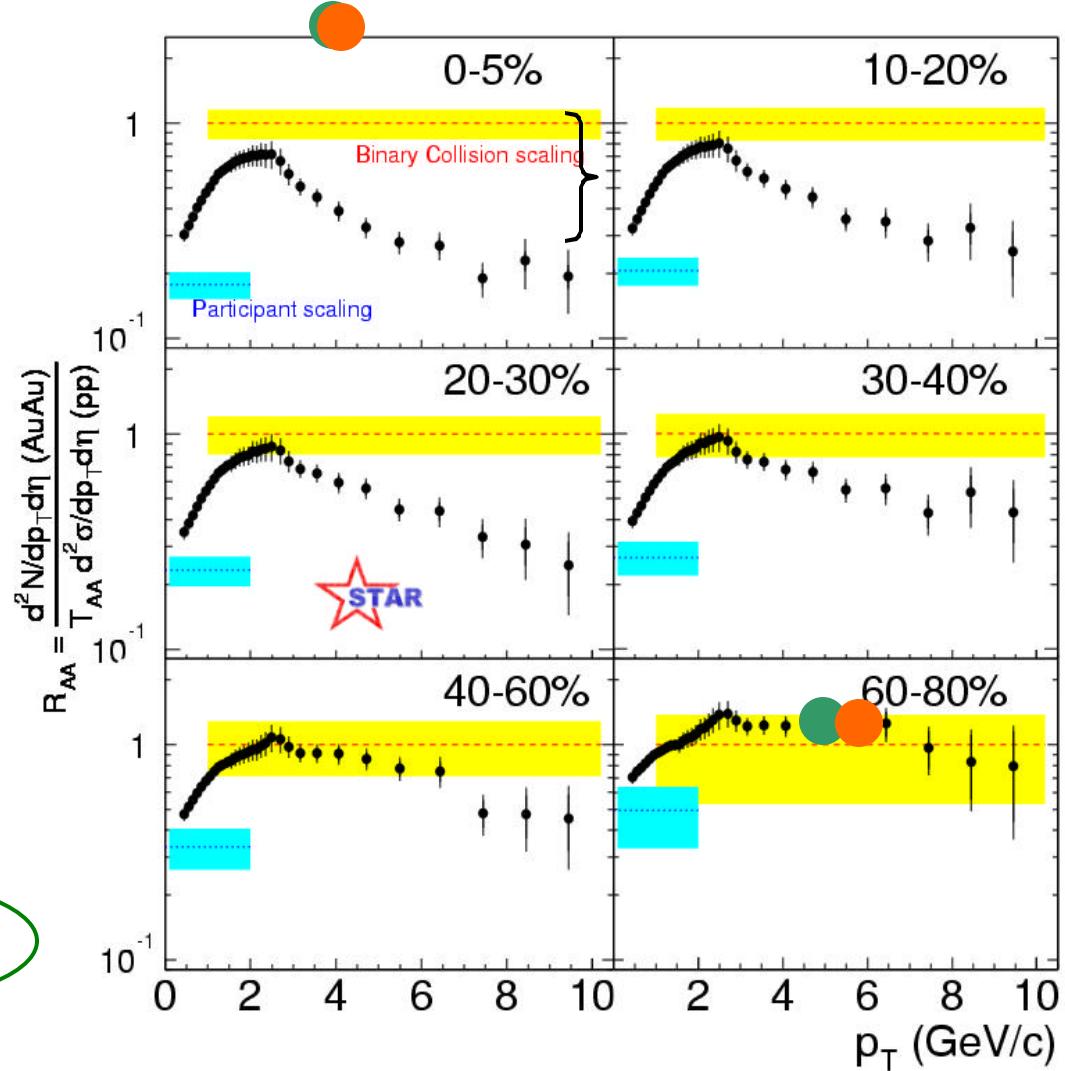
TROP03 06/09/2003

Au+Au relative to p+p

nucl-ex/0305015

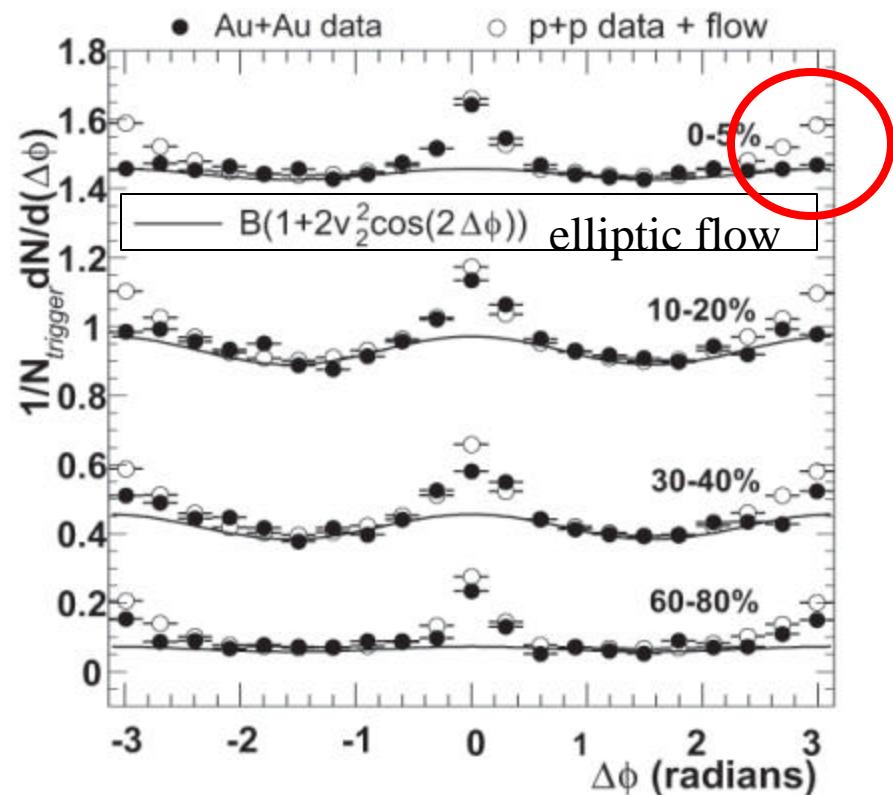
central collisions:
factor ~4-5 suppression

$$R_{AA}(p_T) = \frac{d^2N^{AA}/dp_T d\eta}{T_{AA} d^2\sigma^{NN}/dp_T d\eta}$$

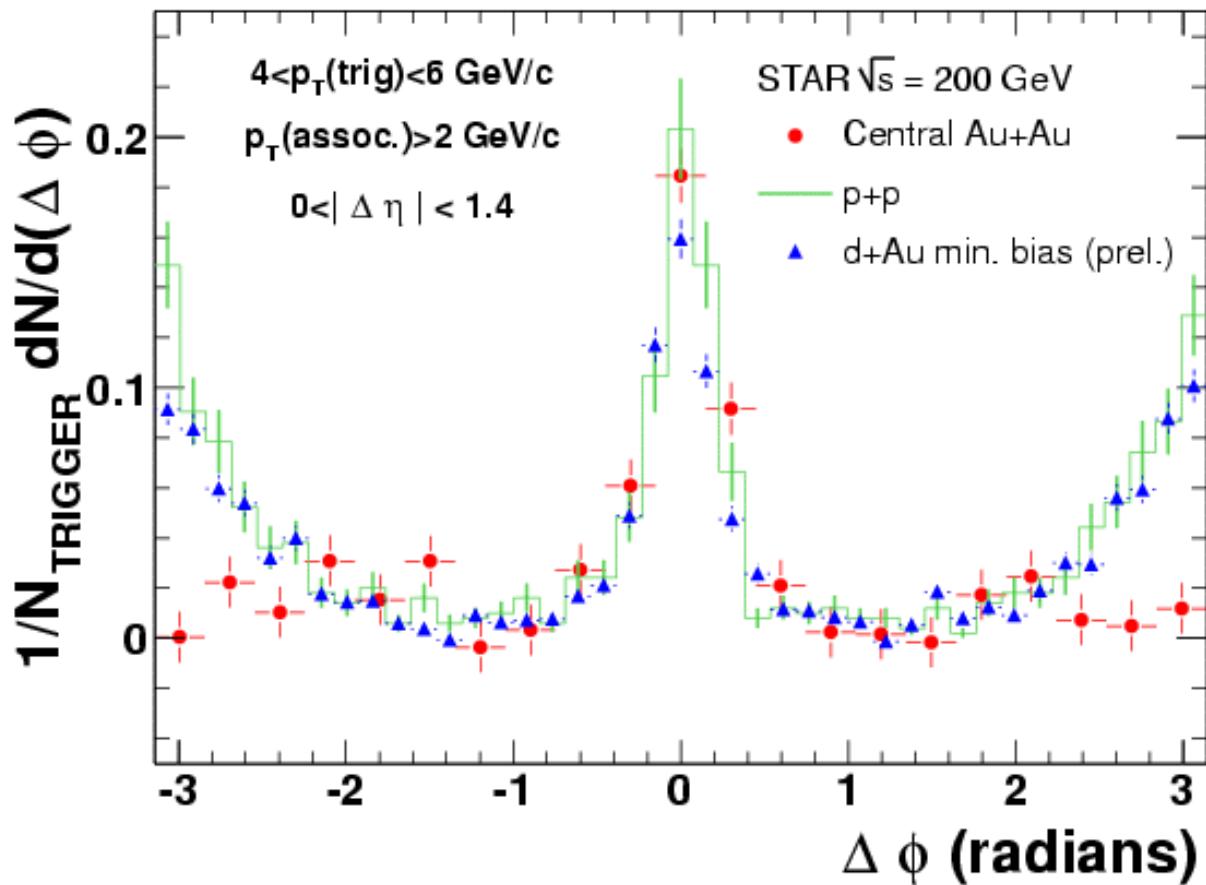


Back-to-back leading jet particles

- Trigger particle ("same side")
 $p_T > 4 \text{ GeV}/c$
- Away-side
 $2 < p_T < 4 \text{ GeV}/c$
- $\Delta\phi$
- Normalize by number
of trigger particles

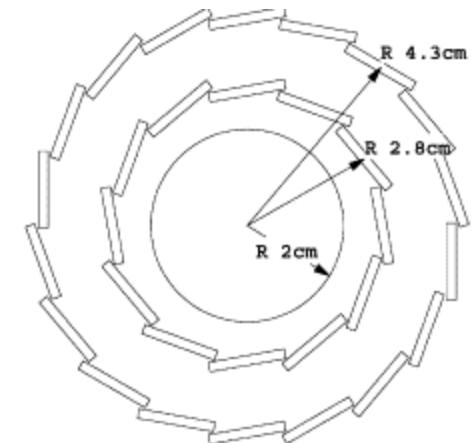
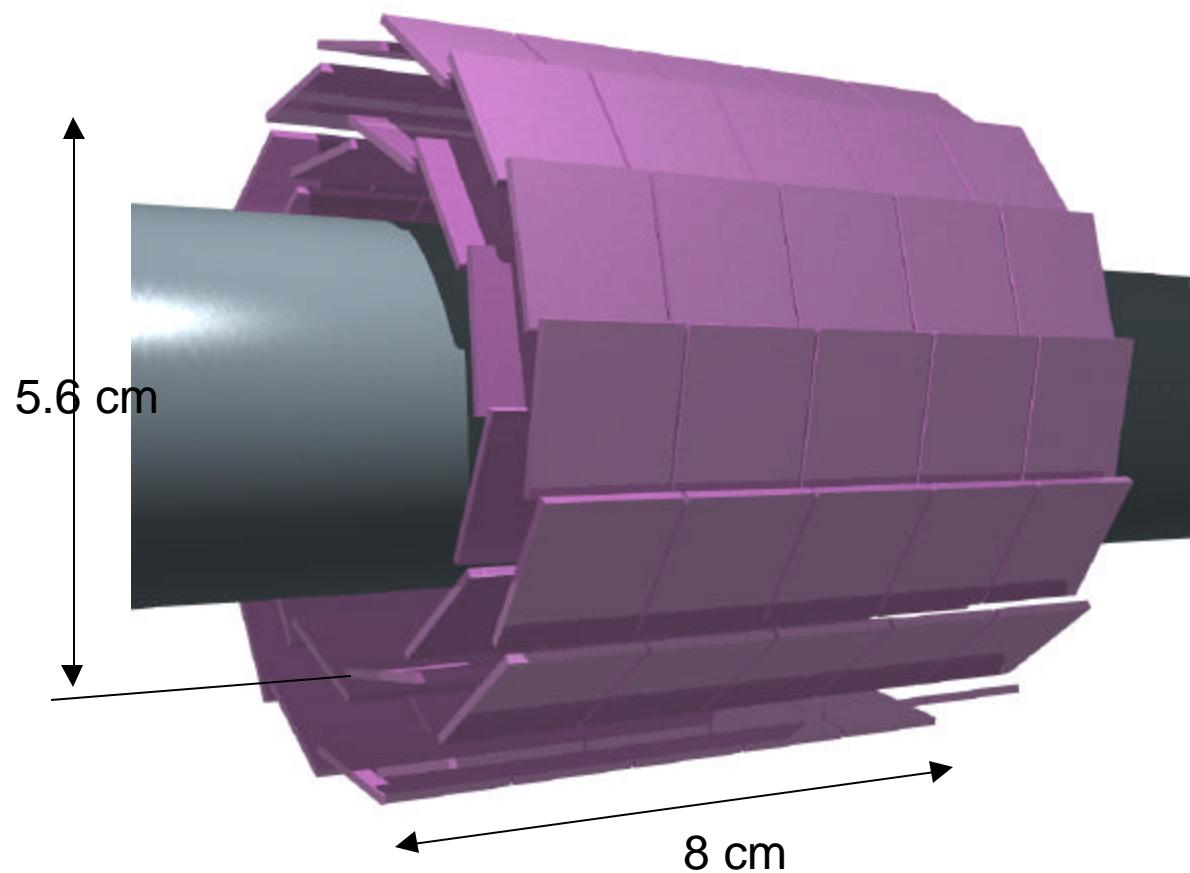


Away-Side Jets in p+p, d+Au, and Au+Au Collisions



STAR Upgrade Pixel Detector

for secondary c,b vertices



- 20 μm square pixels
- $N(\text{pixel}) = 90 \times 10^6$
- $d=40 \mu\text{m}$
(monolithic)
- ADC in pixel
- SRAM in pixel

What did we learn so far ?

- $t < 1 \text{ fm}/c$ temperature: beyond lattice-QCD boundary prediction.
- $t < 1 \text{ fm}/c$ energy density: too high for hadron gas.
- parton-hadron transition (“freeze-out”) short duration.
 $R_{\text{out}}/R_{\text{side}} \sim 1$: There is an explosion
- No indication for 1st order phase transition
(no long-lived slowly-burning mixed quark-hadron soup).
- pQCD models need non-standard extensions,
i.e. high density and high temperature.
- Back-to-back jets disappear.
- Nature recently exhibits surprising macroscopic phenomena
(metallic hydrogen, left-handed materials, clockwise tornadoes).
We will continue to search for microscopic analogons.

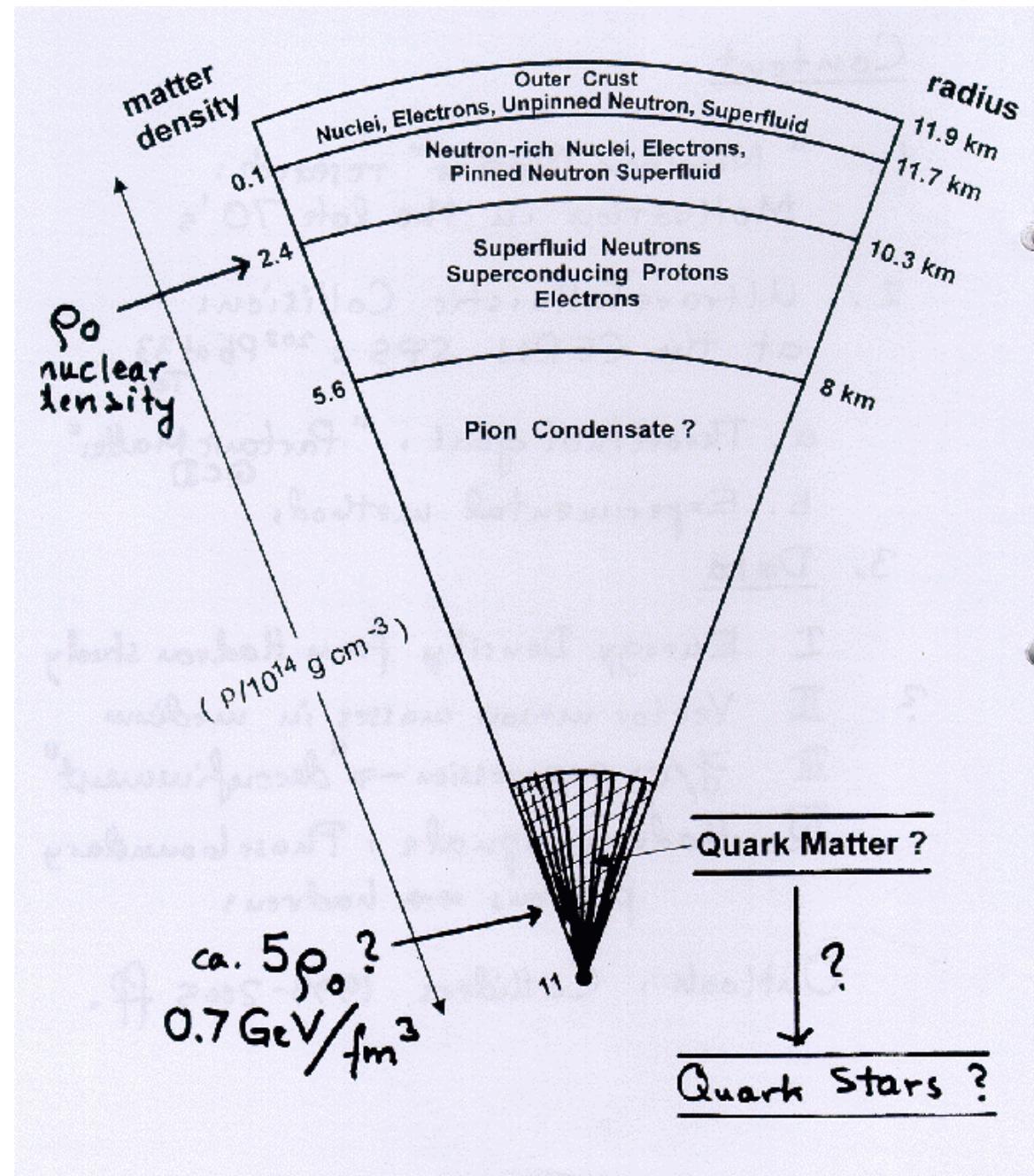
Motto motto,
mou chotto !

What about the big bang ?

- general relativity curvature
 $ds^2 \sim dr^2/(1-k)$
 $k = 2GM/Rc^2$
 $k = 1$ black hole (escape velocity = c)
 $k \sim 10^{-22}$ @ RHIC
→ we are not bending space-time
- universe expands with $v=c$,
 $v=0.55c$ @ RHIC
- big bang = thermal equilibrium
@ RHIC: maybe, not clear yet
but jet dE/dx could mean early equilibrium before suppression
- time scale $1 \text{ fm}/c \sim 10^{-24} \text{ s}$
in the universe: cosmic inflation 10^{-34} s
→ system already had macroscopic size

What about neutron stars ?

- for $R < 1 \text{ km}$ $\rho = 5\rho_0$ expected
we have $15\rho_0$
but energy density much smaller
neutron star $\epsilon = 0.7 \text{ GeV/fm}^3$
- neutron star = hydrostatic equilibrium stability:
outward pressure from incompressibility of nuclear matter
incompressibility = $f(\rho)$
(this is "equation of state")
@ RHIC: outward pressure generates blast wave.



Fit function

Hadron resonance ideal gas

Refs. J.Rafelski PLB(1991)333
J.Sollfrank et al. PRC59(1999)1637

Particle density
of each particle

$$\rho_i = \gamma_s^{|s_i|} \frac{g_i}{2\pi^2} T_{ch}^3 \left(\frac{m_i}{T_{ch}}\right)^2 K_2(m_i/T_{ch}) \lambda_q^{Q_i} \lambda_s^{s_i}$$
$$\lambda_q = \exp(\mu_q/T_{ch}), \quad \lambda_s = \exp(\mu_s/T_{ch})$$

Q_i : 1 for u and d, -1 for \bar{u} and \bar{d}

T_{ch} : Chemical freeze-out temperature

s_i : 1 for s, -1 for \bar{s}

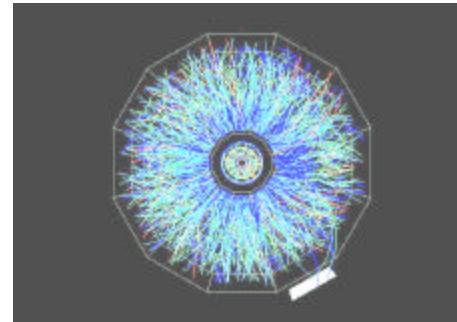
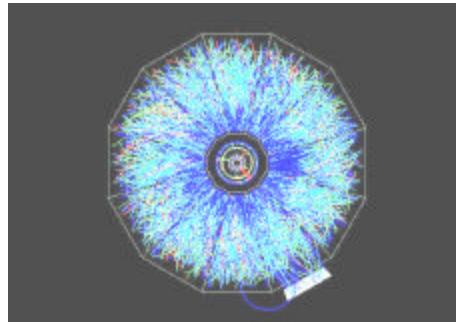
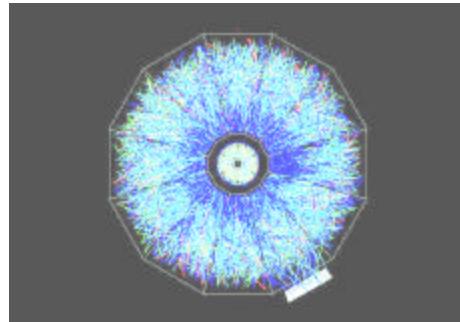
μ_q : light-quark chemical potential

g_i : spin-isospin freedom

μ_s : strangeness chemical potential

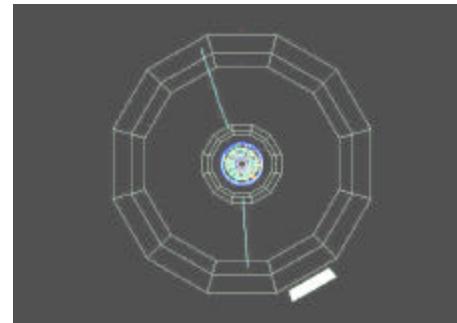
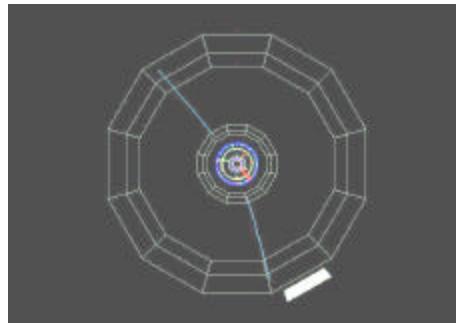
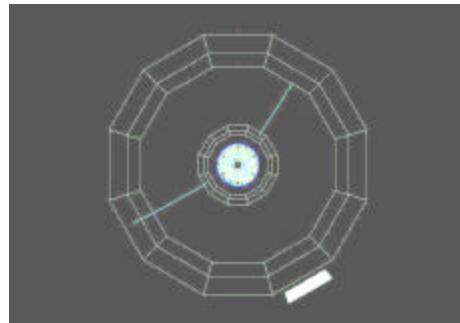
m_i : particle mass

γ_s : strangeness saturation factor



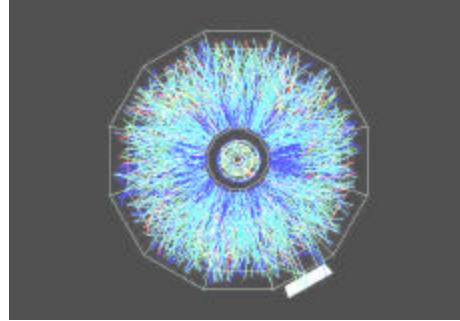
**Upsilon->I+I-
candidates**

**Run 2241022
Event #4699
 $m=10.21 \text{ GeV}$**



**Run 2243027
Event #1645
 $m=9.68 \text{ GeV}$**

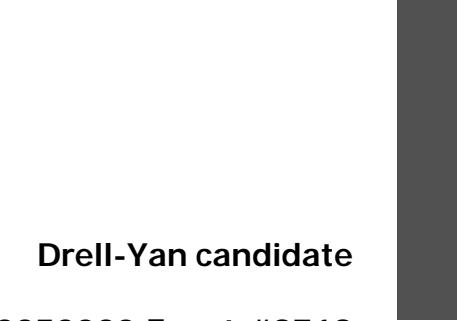
**Run 2244029
Event#1570
 $m=9.25 \text{ GeV}$**



$p^+ p^-$ candidate

**Run 2244032 Event #3420
 $m=9.38 \text{ GeV}$**

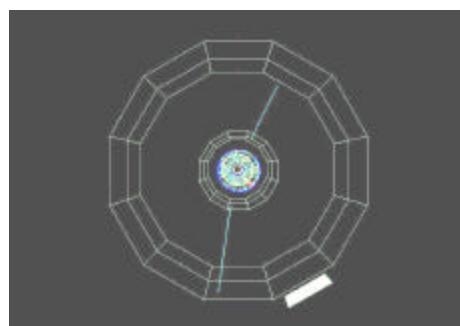
**$dE/dx < 1.4e-5 \text{ keV/cm}$
 $P_{\text{PID}}(\text{lepton}) < 3\%$**



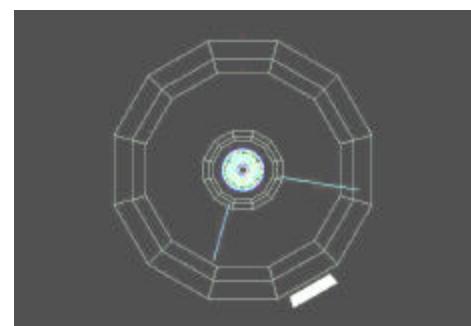
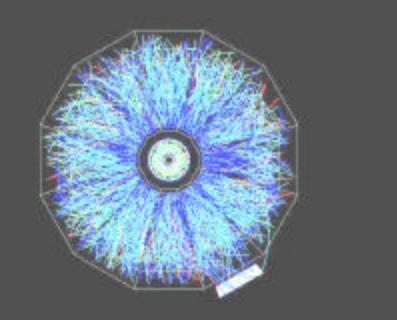
Drell-Yan candidate

**Run 2252020 Event #8718
 $m=10.02 \text{ GeV}$**

$\Delta F = 78^\circ$

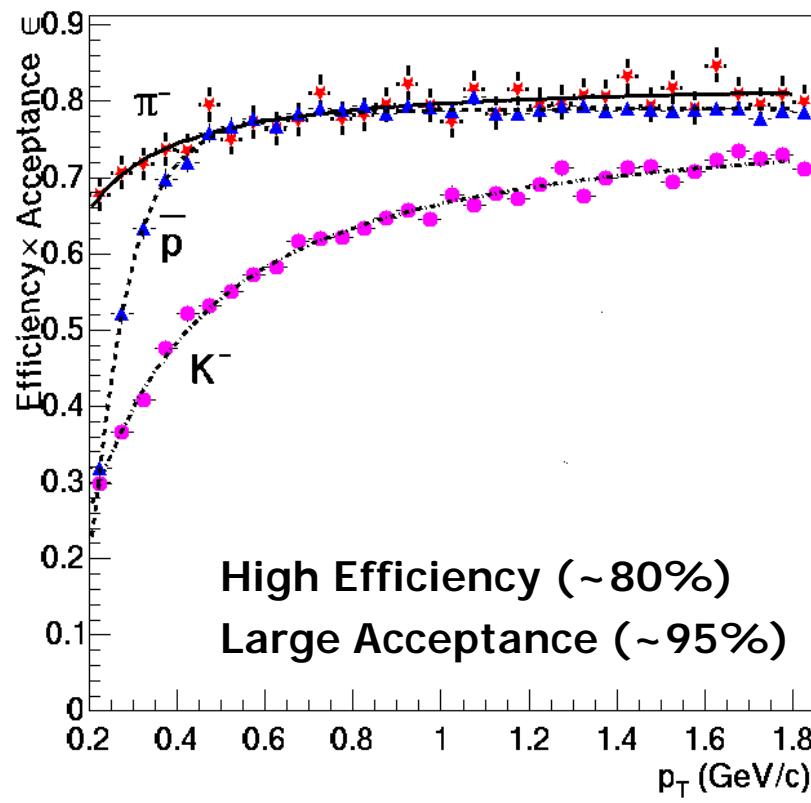


**Au+Au
 $\sqrt{s_{\text{NN}}} = 200 \text{ GeV}$**

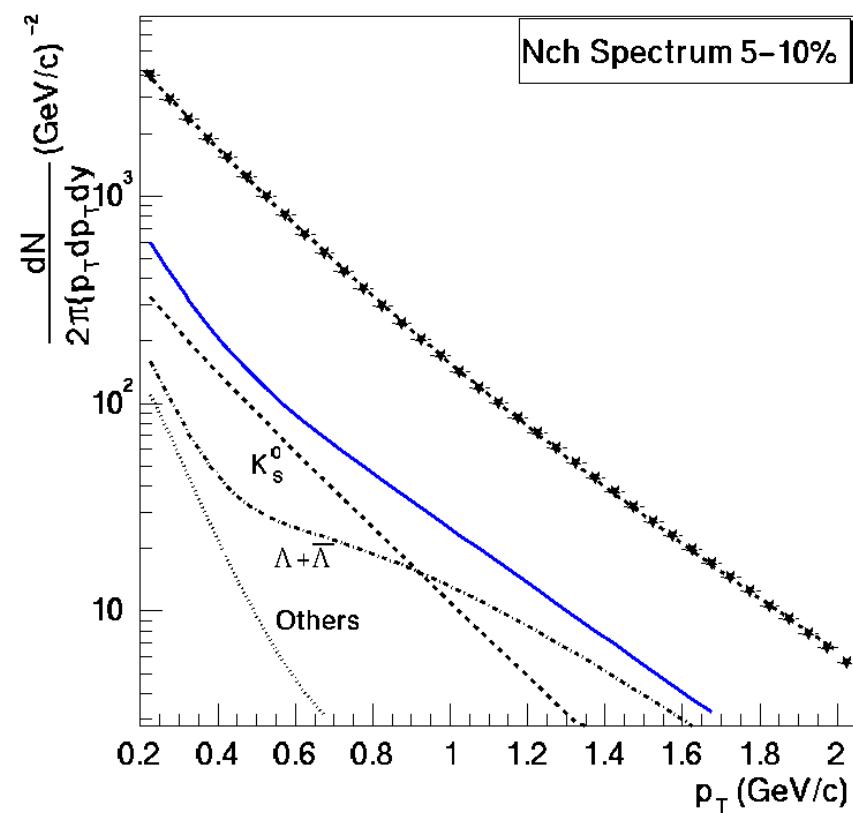


Corrections

Efficiency



Background



The STAR Collaboration

468 Collaborators
10 Countries
49 Institutions
~20 Ph.D. theses
14 PRL
3 PRC

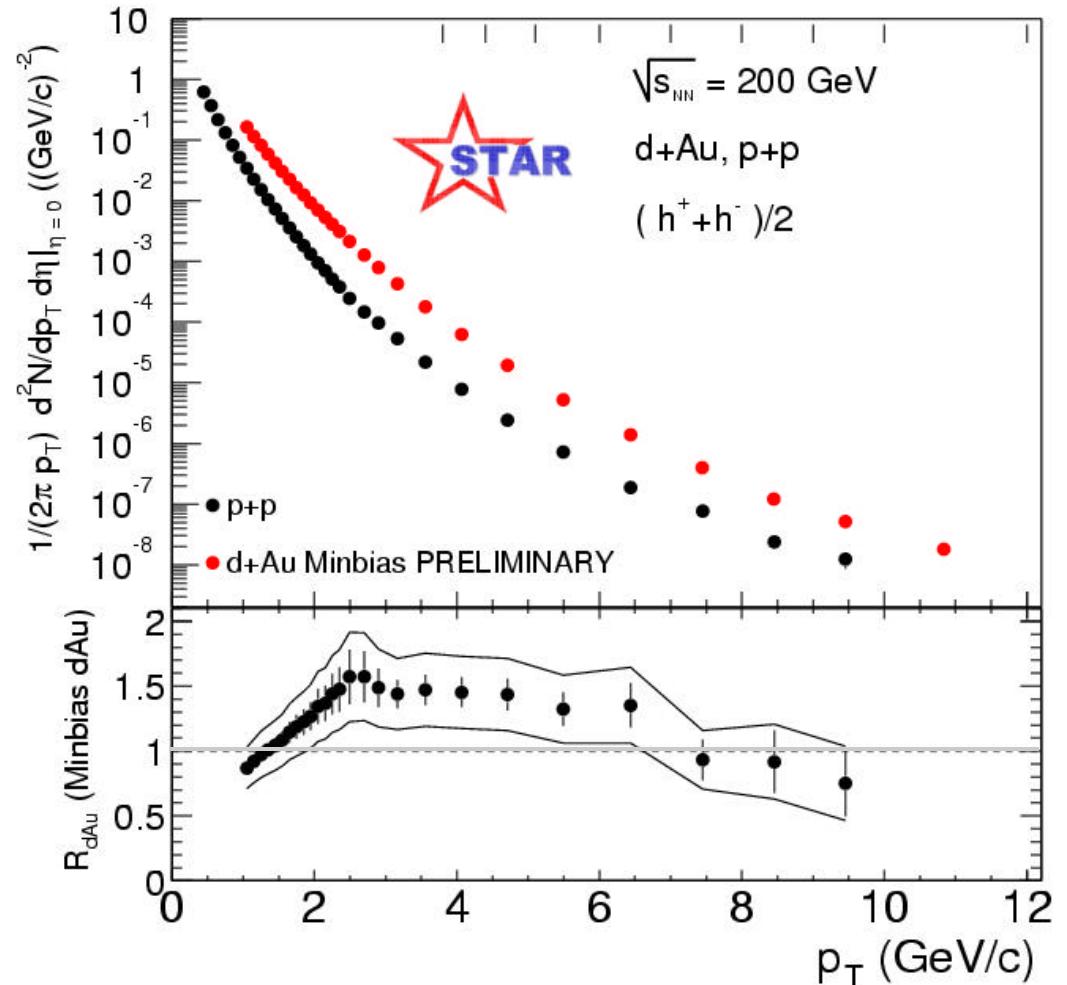
C. Adler¹¹, Z. Ahammed²³, C. Allgower¹², J. Amonett¹⁴, B.D. Anderson¹⁴, M. Anderson⁵, G.S. Averichev⁹, J. Balewski¹², O. Barannikova^{9,25}, L.S. Barnby¹⁴, J. Baudot¹³, S. Bekele²⁰, V.V. Belaga⁹, R. Bellwied²⁰, J. Berger¹¹, H. Bichsel²⁹, L.C. Bland¹², C.O. Blyth³, B.E. Bonner²⁴, R. Bossingham¹⁵, A. Boucham²⁶, A. Brandin¹⁸, R.V. Cadman¹, H. Caines²⁰, M. Calderón de la Barca Sánchez³¹, A. Cardenas²³, J. Carroll¹⁵, J. Castillo²⁶, M. Castro²⁰, D. Cohen³, S. Chaitopadhyay³⁰, M.L. Chen², Y. Chen⁶, S.P. Chernenko⁹, M. Chernyey⁸, A. Chikanian³¹, B. Choi²⁷, W. Christie², J.P. Coffin¹³, L. Conin², T.M. Cormier³⁰, J.G. Cramer²⁹, H.J. Crawford⁴, M. DeMello²⁴, W.S. Deng¹⁴, A.A. Derevitskiov²², L. Didenko², J.E. Draper⁵, V.B. Dunin⁹, J.C. Dunkop³¹, V. Eckardt¹⁶, L.G. Elmendorf⁹, V. Emelianov¹⁸, J. Engelage¹, G. Eppley²⁴, B. Erasmus²⁸, P. Fachini²⁵, V. Faine², E. Finch³¹, Y. Flisyak², D. Flierl¹¹, K.J. Foley², J. Fu¹⁵, N. Gagunashvili¹, J. Gans³¹, L. Gaudichet²⁶, M. Germain¹³, F. Geurts²⁴, V. Ghazikhanian⁶, J. Grabksi²³, O. Grachov³⁰, D. Greiner¹⁵, V. Grigoriev¹⁸, M. Guedon¹³, E. Gushin¹⁸, T.J. Hallman², D. Hardtke¹⁵, J.W. Harris³¹, M. Heffner⁵, S. Heppelmann²¹, T. Herston²³, B. Hippolyte¹³, A. Hirsch², E. Hjort¹⁵, G.W. Hoffmann²⁷, M. Horsley³¹, H.Z. Huang⁶, T.J. Humanic²⁰, H. Hümmer¹⁶, G. Igo⁸, A. Ishihara²⁷, Yu.I. Ivanshin¹⁰, P. Jacobs¹⁵, W.W. Jacobs¹², M. Janik²⁸, I. Johnson¹⁵, P.G. Jones³, E. Judd¹, M. Kaneta¹⁵, M. Kaplan⁷, D. Keane¹⁴, A. Kisiel²⁸, J. Klay⁵, S.R. Klein¹⁵, A. Klyachko¹², A.S. Konstantinov²², L. Kotchenda¹⁸, A.D. Kovalenko⁹, M. Kramer¹⁹, P. Kravtsov¹⁸, K. Krueger¹, C. Kuhn¹³, A.I. Kulikov⁹, G.J. Kunde³¹, C.L. Kunz⁷, R.Kh. Kutuev¹⁰, A.A. Kuznetsov⁹, L. Lakehal-Aya²⁶, J. Lamass-Valverde²⁴, M.A.C. Lamont¹, J.M. Landgraf², S. Lange¹¹, C.P. Lansdell²⁷, B. Laslik³¹, F. Law², A. Lebedev², T. LeCompte¹, R. Lednický⁹, V.M. Leontiev²², M.J. LeVine², Q. Li³⁰, Q. Li¹⁵, S.J. Lindenbaum¹⁹, M.A. Lisa²⁰, T. Ljubick², W.J. Llope²⁴, G. LoCurto¹⁶, H. Long⁶, R.S. Longacre², M. Lopez-Noriega²⁰, W.A. Love², D. Lynn², R. Majka³¹, S. Margolis¹⁴, L. Martin²⁸, J. Marx¹⁵, H.S. Matis¹⁵, Yu.A. Matulenko²², T.S. McShane⁸, F. Meissner¹⁵, Yu. Melnik²², A. Meschanin²², M. Messer¹, M.L. Miller³, Z. Mikheevich⁷, N.G. Minai²⁷, J. Mitchell²⁴, V.A. Moiseenko¹⁰, D. Molz², C.F. Moon²⁷, V. Morozov¹⁵, M.M. de Moura³⁰, M.G. Munhoz²⁵, G.S. Mustafa²⁴, J.M. Nelson³, P. Neyski², V.A. Nikitin¹⁰, L.V. Nogach²², B. Norman¹⁴, S.B. Nurushev²², G. Odyniec¹⁵, A. Ogawa²¹, V. Okorokov¹⁸, M. Oldenburg¹⁶, D. Olson¹⁵, G. Pak²⁰, S.U. Pandey³⁰, Y. Panebratsev⁹, S.Y. Panitkin¹, A.I. Pavlinov³⁰, T. Pawlak²⁸, V. Perevozchikov², W. Peryt²⁸, V.A. Petrov¹⁰, W. Pinganaud²⁶, E. Platner²⁴, J. Pluta²⁸, N. Porile²³, J. Porter²⁷, A.M. Poskanzer¹⁵, E. Potrebenikova⁹, D. Prindle²⁹, C. Pruneau³⁰, S. Radomska²⁸, G. Rai¹⁵, O. Ravel²⁶, R.L. Ray²⁷, S.V. Razin^{9,12}, D. Reichhold⁸, J.G. Reid¹⁹, F. Retiere⁵, A. Ridiger¹⁸, H.G. Ritter¹⁵, J.B. Roberts²⁴, O.V. Rogachevskiy⁶, J.L. Romero⁵, C. Roy²⁶, D. Russ⁷, V. Rykov³⁰, I. Sakrejda¹⁵, J. Sandweiss³¹, A.C. Saulys², I. Savin¹⁰, J. Schambach²⁷, R.P. Scharenberg²³, K. Schweda¹⁵, N. Schmitz¹⁶, L.S. Schroeder¹⁵, A. Schiuttau¹⁶, J. Seger⁸, D. Seliverstov¹, P. Seyboth¹⁶, E. Shahaliev⁹, K.E. Shestermanov²², S.S. Shnianshikov⁹, V.S. Shvetsov¹⁰, G. Skoro⁹, N. Smirnov³¹, R. Snellings¹⁵, J. Sowinski¹², H.M. Spinka¹, B. Srivastava²¹, E.J. Stephenson¹², R. Stock¹¹, A. Stolpovsky³⁰, M. Strikhanov¹⁸, B. String fellow²³, H. Stroebel¹¹, C. Struck¹¹, A.A.P. Suakko¹⁰, E. Sugarbaker²⁰, C. Suire¹³, M. Sumbera⁹, T.J.M. Symons¹⁵, A. Szanto de Toledo²⁸, P. Szarwas²⁸, J. Takahashi²⁸, A.H. Tang¹⁴, J.H. Thomas¹⁵, V. Tikhomirov¹⁸, T.A. Trainor²⁹, S. Trentalange⁸, M. Tokarev⁹, M.B. Tonjeas⁷, V. Tlumov¹⁸, O. Tsai⁶, K. Turner², T. Ulrich², D.G. Underwood⁴, G. Van Buren², A.M. VanderMolen¹⁷, A. Vanyashin¹⁶, I.M. Vasilevskiy¹⁰, A.N. Vasiliev²⁷, S.E. Vigdor¹², S.A. Voloshin³⁰, F. Wang²⁷, H. Ward²⁷, J.W. Watson¹⁴, R. Wels²⁰, T. Wenau², G.D. Westfall¹⁷, C. Whitten Jr.⁶, H. Wieman¹⁵, R. Wilson²⁰, S.W. Wissink⁹, R. Wit¹⁴, N. Xu¹⁵, Z. Xu³¹, A.E. Yakutin²², E. Yamamoto⁶, J. Yang⁶, P. Yepes²⁴, A. Yokosawa¹, V.I. Yurevich⁹, Y.V. Zanevski⁹, I. Zborovsky⁹, W.M. Zhang¹⁴, R. Zoukarneev¹⁰, A.N. Zubarev⁹

d+Au: minbias charged hadron yield

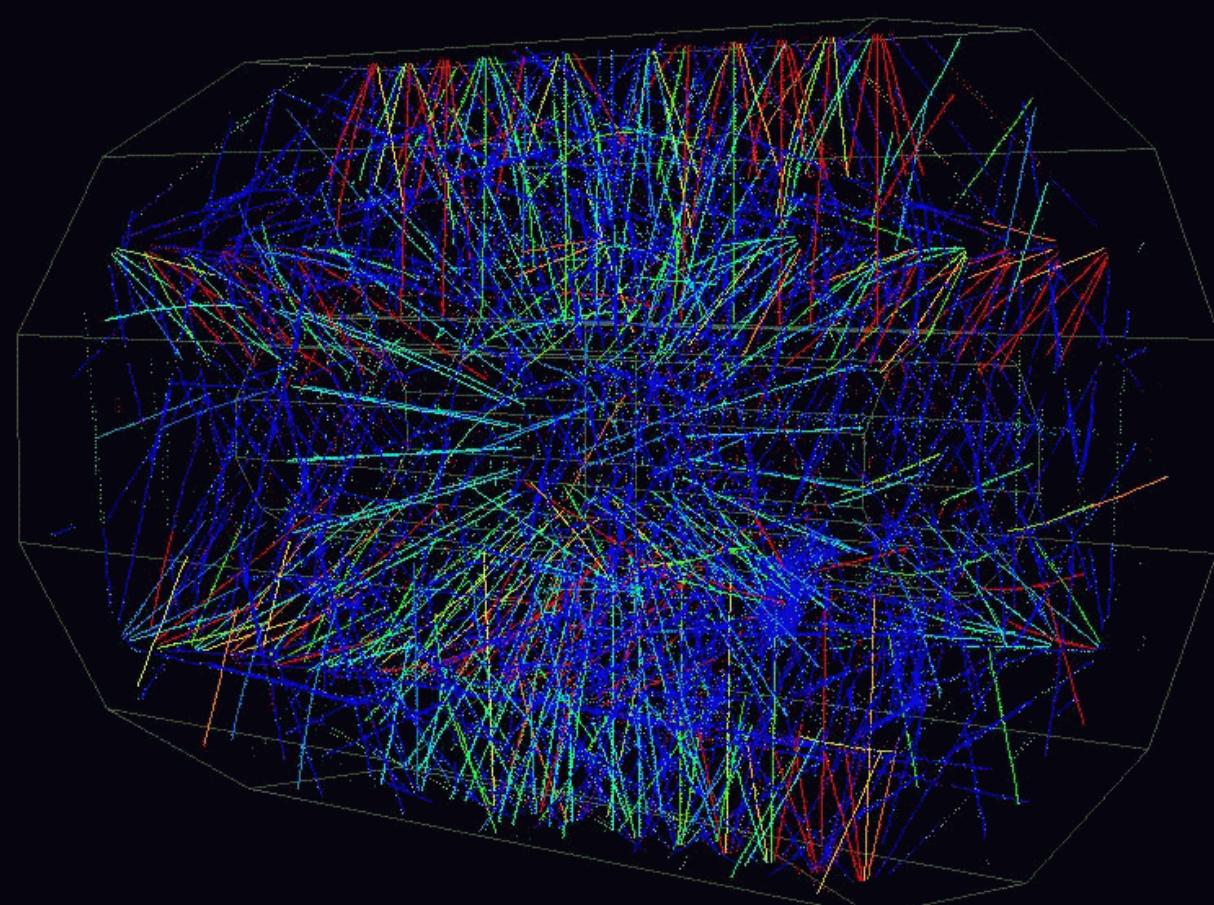
$$R_{dAu} = \frac{d^2N^{dAu}/dp_T d}{T_{dAu} d^2 p_{pp}/dp_T d}$$

$$T_{dAu} = \frac{2 \times 197}{2.2 b}$$

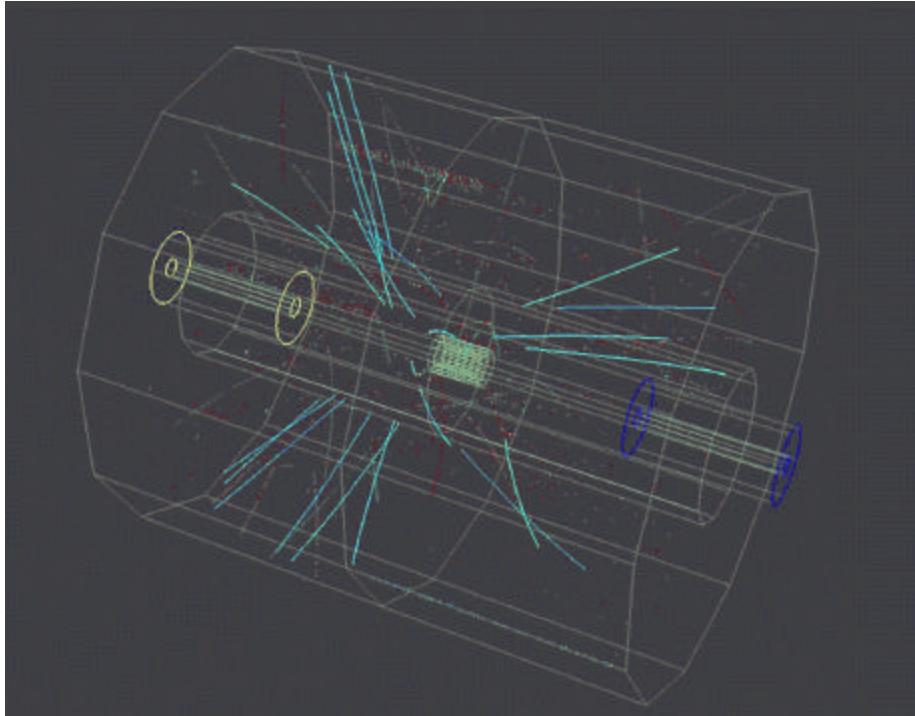
$R_{AA} > 1$
 charged hadron yield
enhanced in d+Au
 Cronin effect
 (expected for "cold"
 nuclear matter)



Laser for TPC Calibration

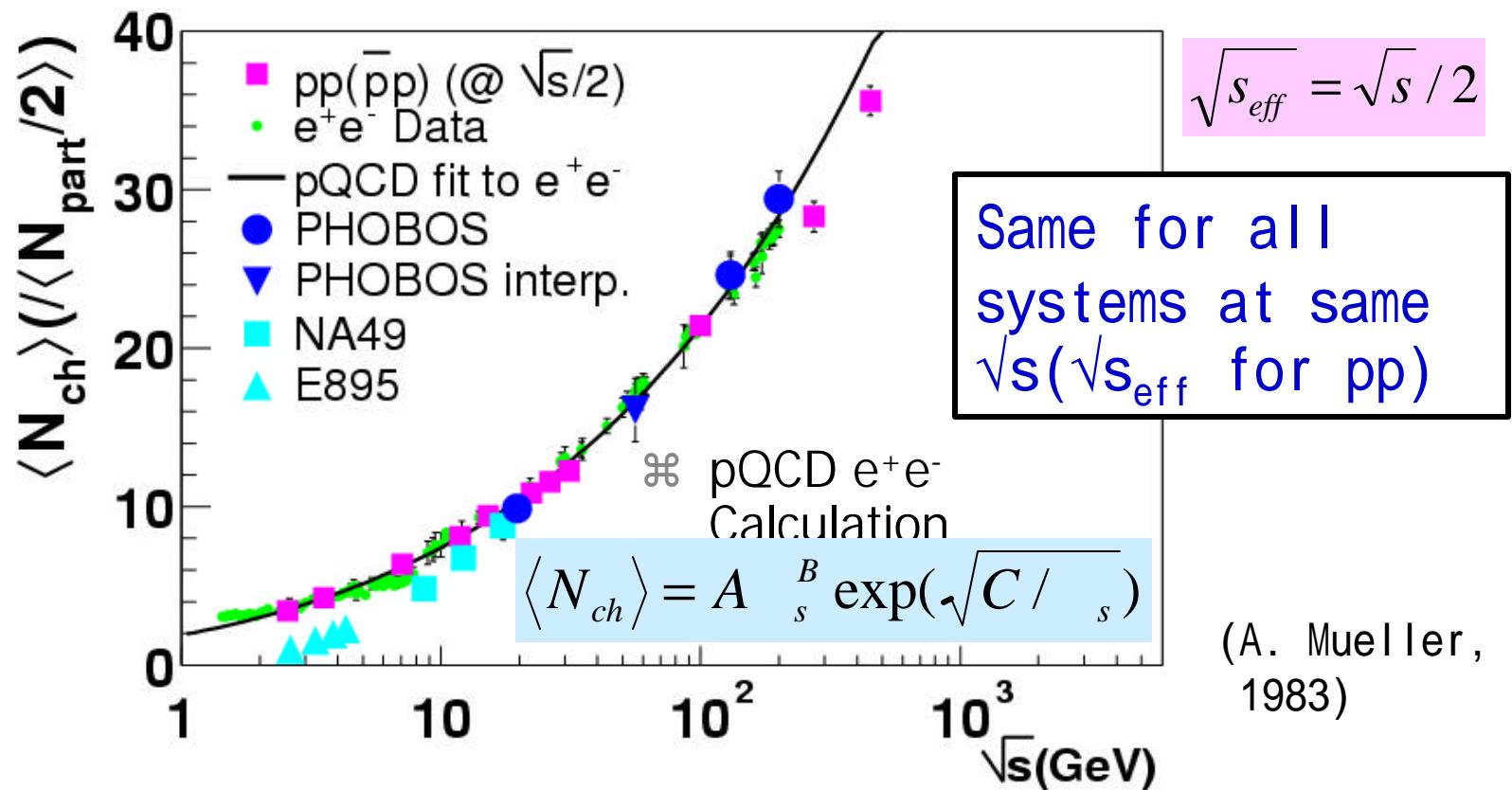


3-Jet Event (pp , $\sqrt{s}=200 \text{ GeV}$)



3 jet (ggg or qq γ or Upsilon decay)
all tracks $p_T > 0.5 \text{ GeV}/c$

$N_{\text{charged}}(\ddot{\sigma}_{NN})$ – Universality of Total Multiplicity?



Total charged particle multiplicity / participant pair